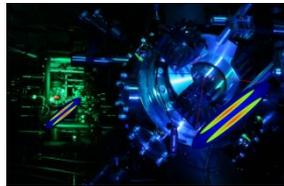
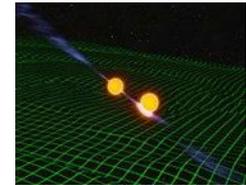


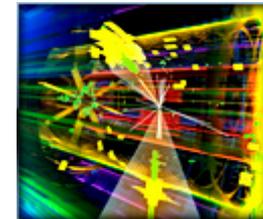


# News from NSF MPS & PHY

Denise Caldwell



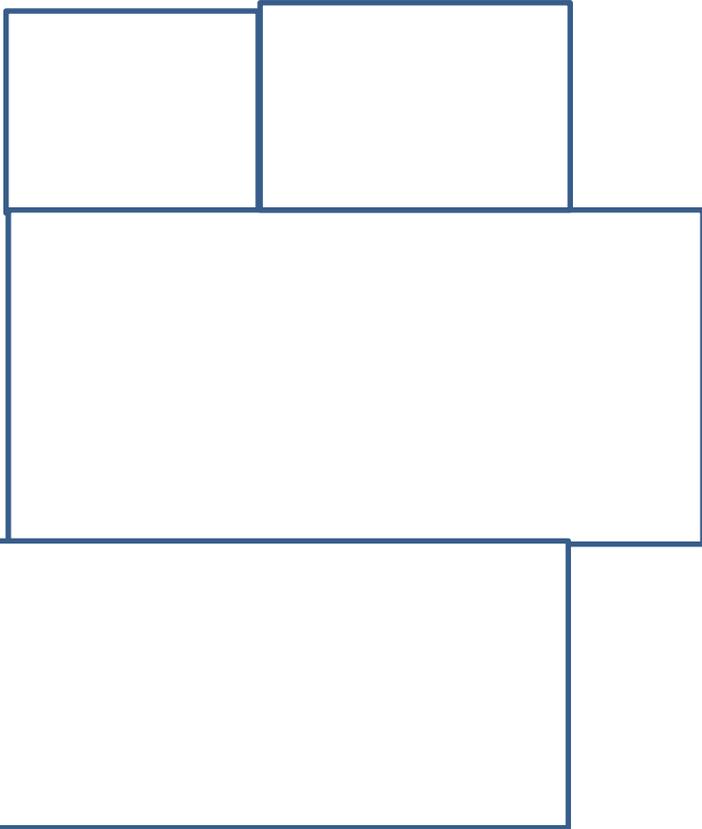
Division Director  
Division of Physics





# First Observation of Binary Neutron Star Merger

## LIGO + VIRGO + 60 Observatories



space-time  
ripples



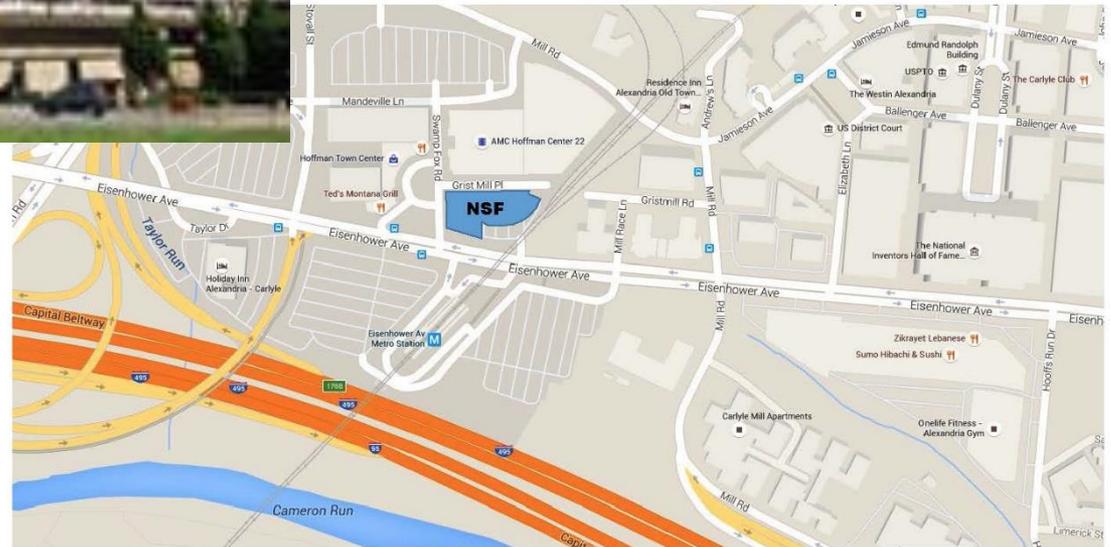
matter  
distortions



# NSF Has a New Home



2415 Eisenhower Avenue  
Alexandria, VA 22314





# Mathematical and Physical Sciences



# New Assistant Director Designated

Welcome (January 2)

**Dr. Anne Kinney**



BS Degree in Astronomy & Physics; PhD in Astrophysics

More than 30 years of leadership and management experience in the astronomical community

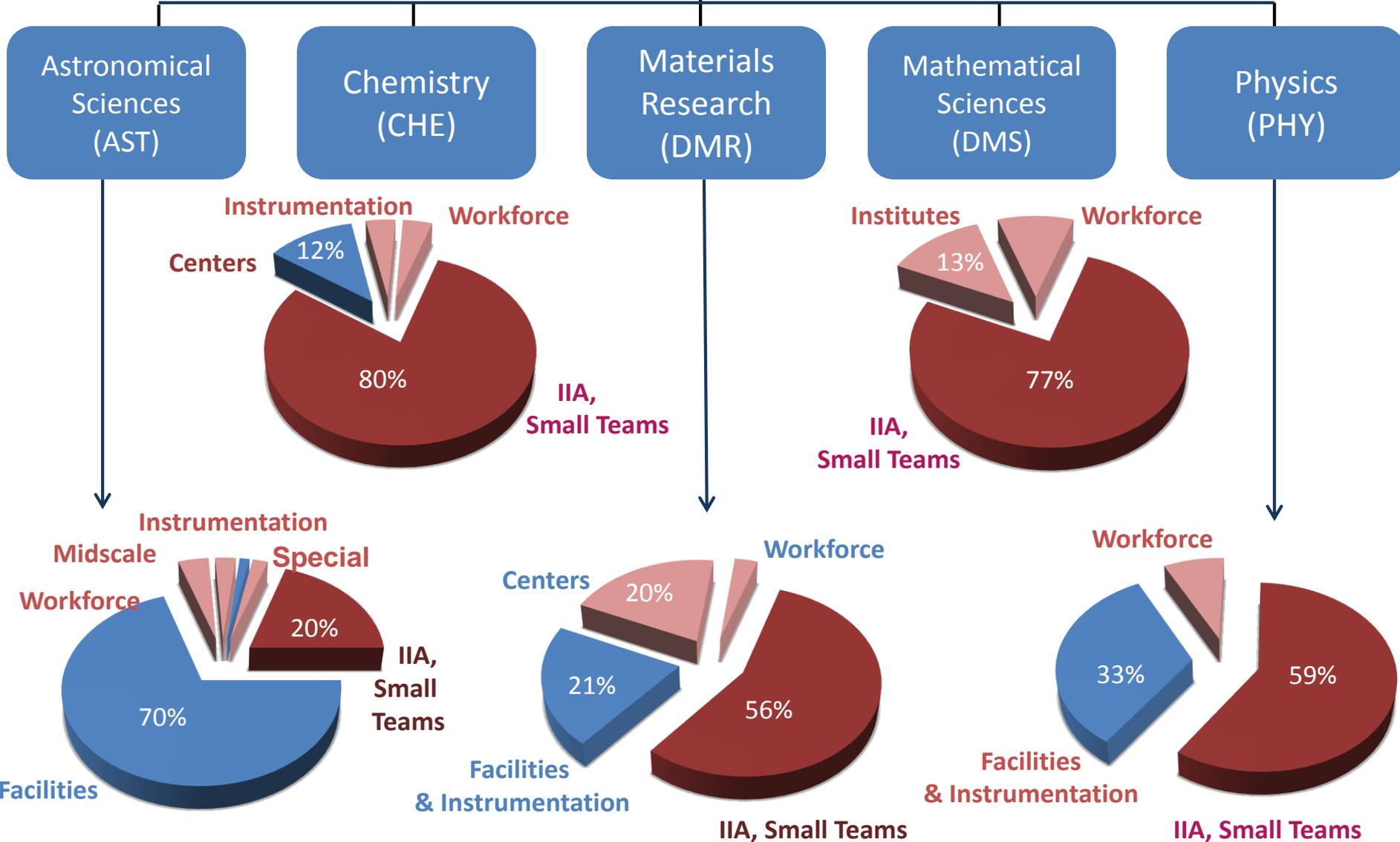
Director of Universe Division at NASA – Oversaw missions including Hubble Space Telescope, Spitzer Space Telescope, WMAP, Galaxy Evolution Explorer

NASA Goddard Space Flight Center – Most recently As Director of Solar System Exploration Division

Since 2015, chief scientist at the W. M. Keck Observatory



# Mathematical and Physical Sciences Directorate





# MPS in FY 2018 Request

## MPS Funding

(Dollars in Millions)

	FY 2016 Actual	FY 2017 Estimate	FY 2018 Request	Change Over FY 2016 Actual	
				Amount	Percent
Astronomical Sciences (AST)	\$246.63	246.16	\$221.15	-\$25.48	-10.3%
Chemistry (CHE)	246.52	245.74	221.05	-25.47	-10.3%
Materials Research (DMR)	309.88	314.31	282.87	-27.01	-8.7%
Mathematical Sciences (DMS)	233.95	233.51	209.78	-24.17	-10.3%
Physics (PHY)	276.91	281.38	253.30	-23.61	-8.5%
Office of Multidisciplinary Activities (OMA)	34.89	34.93	31.28	-3.61	-10.3%
<b>Total</b>	<b>\$1,348.78</b>	<b>\$1,356.03</b>	<b>\$1,219.43</b>	<b>-\$129.35</b>	<b>-9.6%</b>

- Smaller DMR and PHY decreases in FY 2017 caused by hosting of new NSF Science and Technology Centers
- Construction of DKIST and LSST are in a separate budget line, fully funded in request



# FY 2018 Appropriation Status

- Congress has passed a Continuing Resolution through December 8, at approximately FY 2017 funding levels
- Note: Budget Control Act (aka “sequester”) still in effect
- Appropriations Subcommittee Bills (in \$M):

Line	FY16 Actual	FY17 Enacted	FY 18 Request	House	Senate
NSF	7,494	7,472	6,653	7,339	7,311
R&RA	5,998	6,034	5,362	6,034	5,918
EHR	884	880	761	880	862
MREFC	242	209	183	78	183
AOAM	351	330	329	329	329
NSB	4	4	4	4	4
OIG	15	15	15	15	15



## Division of Physics



# FY 2018 PHY Budget Request

PHY Funding					
(Dollars in Millions)					
	FY 2016 Actual	FY 2017 (TBD)	FY 2018 Request	Change Over	
				FY 2016 Actual Amount	Percent
Total	\$276.91	-	\$253.30	-\$23.61	-8.5%
Research	174.12	-	152.09	-22.03	-12.7%
CAREER	8.12	-	7.30	-0.82	-10.1%
STC: Center for Bright Beams (CBB)	-	-	5.00	5.00	N/A
Education	5.40	-	4.80	-0.60	-11.1%
Infrastructure	97.39	-	96.41	-0.98	-1.0%
IceCube	3.48	-	3.50	0.02	0.6%
Large Hadron Collider (LHC)	20.00	-	16.00	-4.00	-20.0%
Laser Interferometer Gravitational Wave Observatory (LIGO)	39.43	-	39.43	-	-
Nat'l Superconducting Cyclotron Lab. (NSCL)	24.00	-	23.00	-1.00	-4.2%
Midscale Research Infrastructure	10.48	-	8.18	-2.30	-21.9%
Pre-construction planning:	-	-	6.30	6.30	N/A
High-Luminosity LHC Upgrade Planning	-	-	6.30	6.30	N/A



# Particle Physics in Physics Division

## EPP – Experimental Elementary Particle Physics

(Mostly Accelerator-Based users, incl. LHC)

Program Director: Saul Gonzalez

Program Director: Jim Shank (sadly leaving in January ☹️)

Program Director: Randy Ruchti (happily joining in January 😊)

## PA – Experimental Particle Astrophysics

(Underground Physics, Cosmic Physics, IceCube users)

Program Director: Jim Whitmore

Program Director: Jean Cottam-Allen

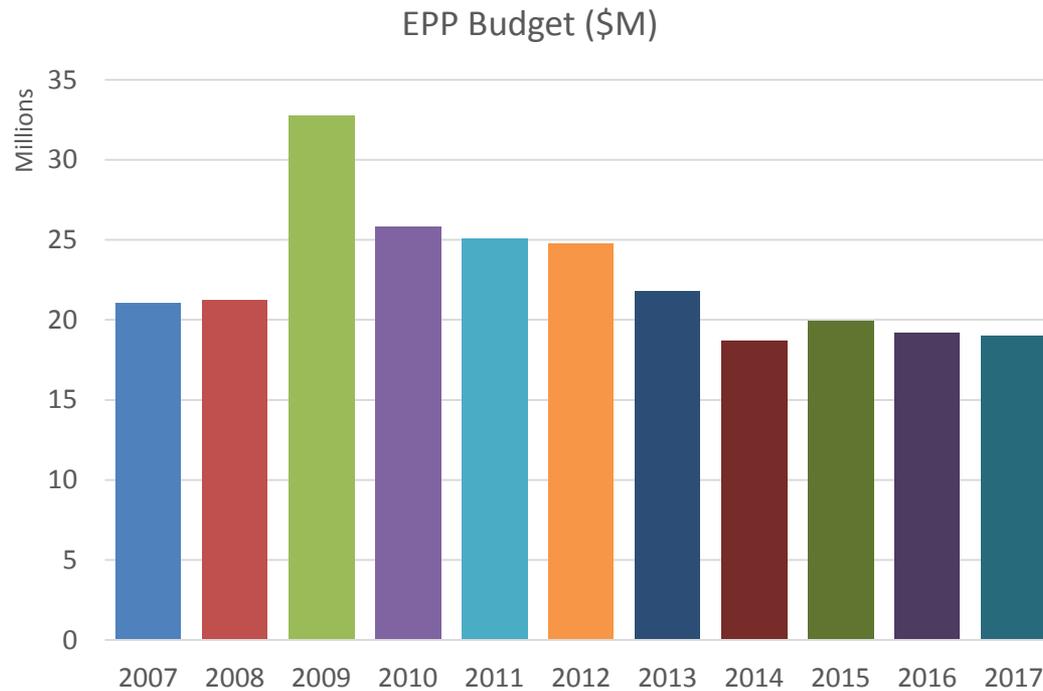
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## Theory – Elementary Particle Physics & Astrophysics and Cosmology

Program Director: Keith Dienes



# EPP Budget by FY



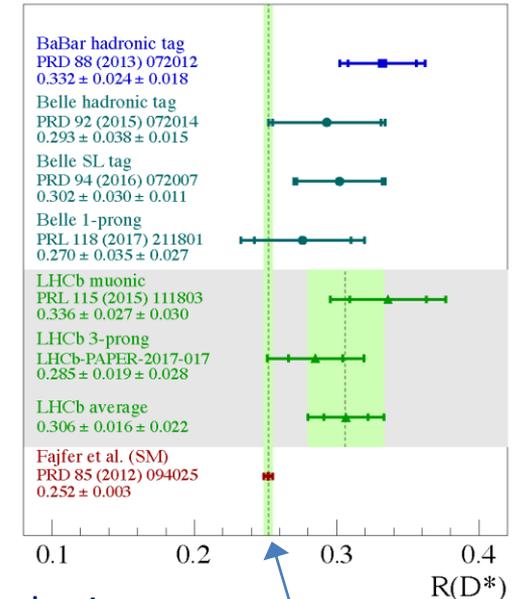
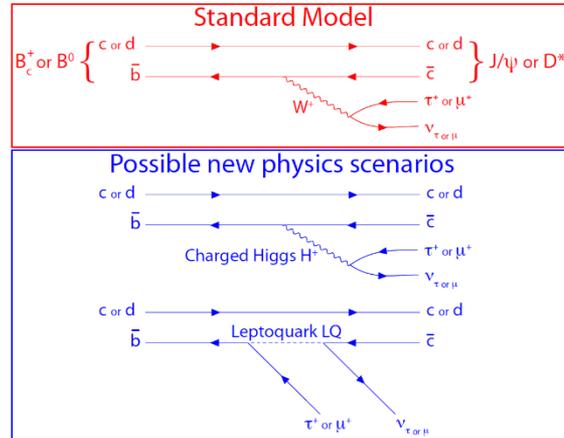
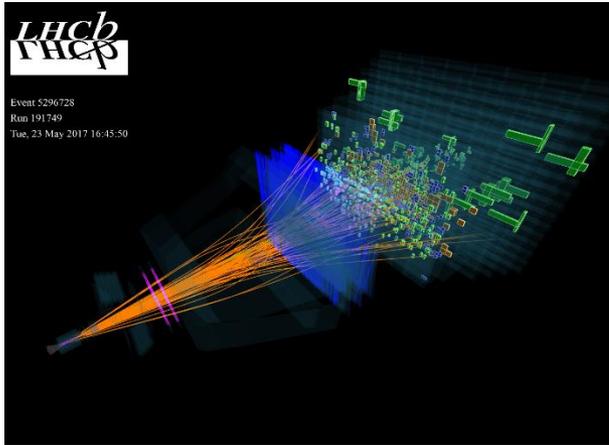


# EPP 2017

- Total funding: \$18,972,899
- Commitments: \$13,308,386
- FY17 Renewals: \$5,664,513
- Number of Proposals: 38
- Number funded: 7, All at lower than the requested amount
- 1 new CAREER Award:
  - Verena Martinez Outschoorn, UIUC.



# LHCb experiment: lepton universality



LHCb is the smallest of 4 experiments at the LHC, focused on finding new physics through precision measurements of  $b$  quarks. NSF provides support to 5 of 72 research groups in a collaboration of 820 physicists.

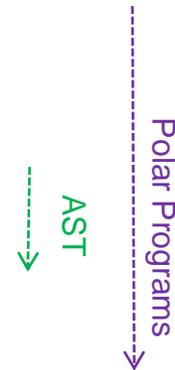
- Ongoing Lepton Universality test probes physics beyond the Standard Model
- Ratio of  $B_c \rightarrow J/\psi \mu \nu_\mu$  rate and  $B_c \rightarrow J/\psi \tau \nu_\tau$  rate is very sensitive to new physics, such as charged Higgs or leptoquarks
- NSF groups made key contributions to this analysis showing possible deviation from the Standard Model

Standard Model



# PA Program Scope - Supported Projects

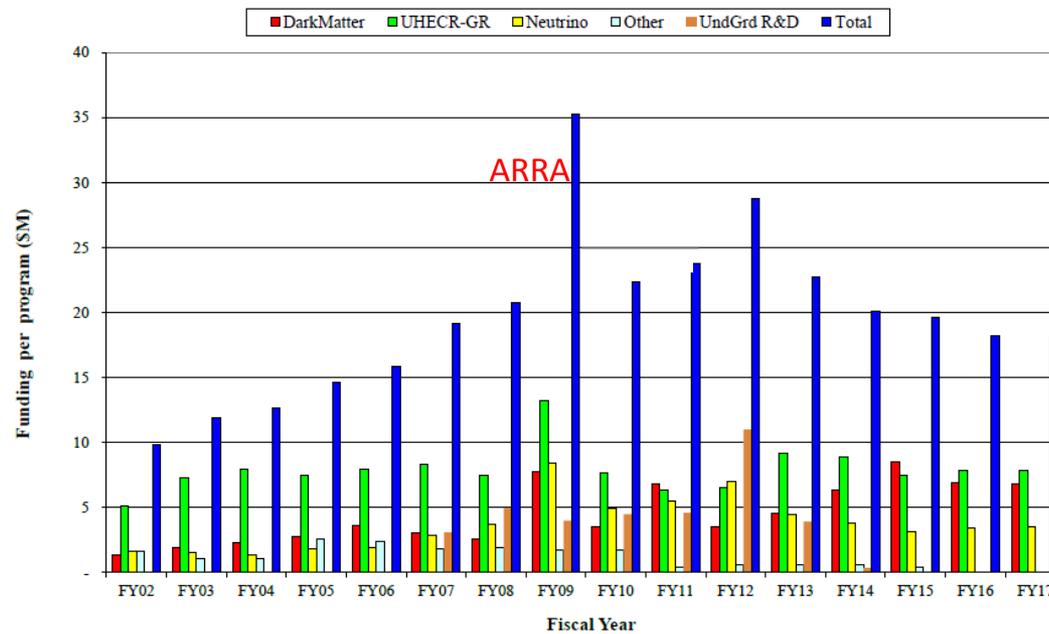
- Direct Dark Matter Detection – WIMP and non-WIMP experiments  
SuperCDMS at SNOLAB, XENON100/1T, LUX, DArkSide-50, PICO, DRIFT, DM-Ice, SABRE, DAMIC, HAYSTAC (ADMX-HF), ALPS2 and Light mass DM experiments
- Indirect Dark Matter Detection  
VERITAS, HAWC, IceCube
- Cosmic Ray, Gamma Ray, and UHE Neutrino Observatories  
IceCube, VERITAS, HAWC, Auger, Telescope Array, CTA, ARA, ARIANNA
- Cosmic Microwave Background  
SPT and BICEP
- Neutrino Properties  
Double Chooz, Project 8, IceCube, IsoDAR, CHANDLER
- Solar, SuperNova and Geo-Neutrinos  
Borexino, SNEWS
- Detector R&D  
NaI/CsI, LiSc/QuDots





# PA Program Funding FY2002-2017

Yearly Funding for overall PA Program (\$M)





# HAWC Results: Constraints on Origin of Positron Flux

“Extended gamma-ray sources around pulsars constrain the origin of the positron flux at Earth,” *Science*, November 2017

HAWC measured extended high-energy gamma ray emission around two nearby pulsars and constrained the diffusion rates of the accelerated high-energy leptons. The low diffusion rate from pulsars suggests a more exotic origin for the high flux of positron emission observed on Earth.

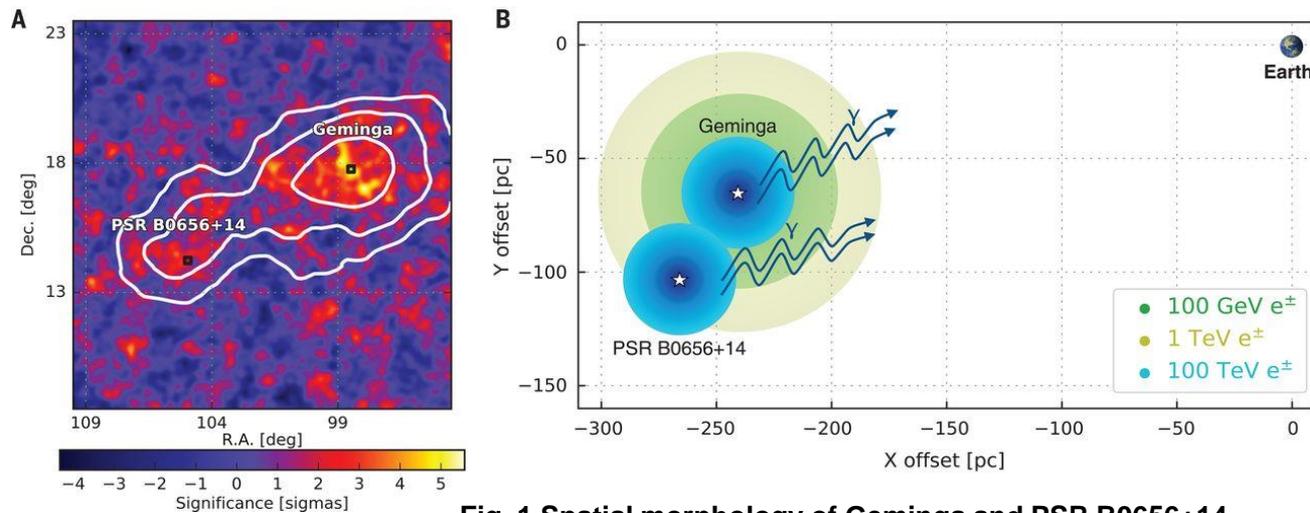


Fig. 1 Spatial morphology of Geminga and PSR B0656+14.



# IceCube Results: Measurement of Neutrino interaction Cross-Sections

“Measurement of the multi-TeV neutrino interaction cross-section with IceCube using Earth absorption,” Nature, November 2017

IceCube reports measurements of neutrino absorption by the Earth using observations of upward-moving neutrino-induced muons. They determined the neutrino–nucleon interaction cross-section for neutrinos at energies of 6.3–980 TeV, which is more than an order of magnitude higher than previous measurements.

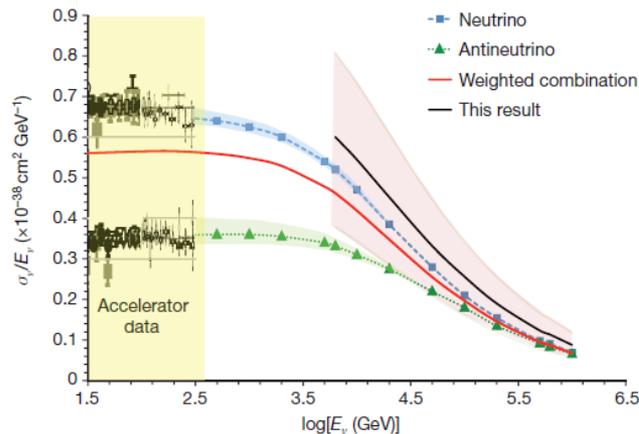


Fig. 1 Neutrino Cross Section Measurement

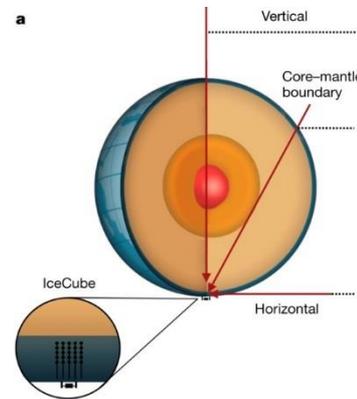


Fig. 2a Neutrino Absorption in the Earth



# Theory

A vibrant, intellectually diverse Theory program is vital to the success of the entire Particle Physics mission. We capitalize on the talents and creativity of the Theory community by supporting the best, most cutting-edge investigator-driven research in two programs:

- Theoretical High-Energy Physics
- Theoretical Particle Astrophysics and Cosmology

These two theory programs interface regularly with many other programs at NSF (EPP, PA, Gravity Theory, Nuclear Theory, Astronomy, Materials Research, Mathematical Sciences, etc.) We also coordinate, as needed, with DOE.

Approximately 110 separate active grants supporting ~180 PIs; ~30 large university groups.

Supporting individuals, RUI's, and special facilities or initiatives (Aspen Center for Physics, TASI summer school, LHC Theory Initiative, etc.)



# Theory Trends

- FY15-17: three-year absorption of string-theoretic portion of former Mathematical Physics program. Now nearly complete.
- FY16: NSF renews Aspen Center for Physics grant for next five years, expands support and scope into Atomic Physics
- Numbers of proposals received is currently twice what it was only 3-4 years ago.
- Increasing numbers of RUI proposals, particularly in FY17.
- One major challenge affecting Theory is the entrance of non-traditional (private philanthropic) funding sources. NSF has developed new procedures for evaluating overlapping sources of funding and introducing such evaluations into the proposal review process. (See NSF 17-561)

	FY 2015	FY 2016	FY 2017
THY Budget	\$13.7 million	\$13.2 million	\$13.4 million
Proposals receiving awards	28	30	26
CAREER	2	1	2



# NSF Beyond the Disciplinary Programs



# Centers and Institutes

## Physics Frontiers Centers – FY 2017 Competition Concluded

Kavli Institute for Cosmological Physics – U Chicago

Center for Ultracold Atoms – MIT/Harvard

PFC@JILA – UC Boulder

Kavli Institute for Theoretical Physics – UCSB

Center for Theoretical Biological Physics – Rice

Joint Institute for Nuclear Astrophysics – Michigan State

PFC@Joint Quantum Institute – U Maryland

Center for the Physics of Living Cells – UIUC

Institute for Quantum Information and Matter – CalTech

North America Nanohertz Observatory for Gravitational Waves – UW Milwaukee

Center for the Physics of Biological Function – Princeton/CUNY (New in FY 2017)

## Centers and Institutes (Science and Technology Center)

Center for Bright Beams – Cornell



# Instrumentation

Major Research Instrumentation (MRI) – up to \$4M – [NSF 18-513](#)

Mid-Scale (PHY Only) - \$4M - Realistically \$10-15M TPC; Apply to PHY program

ATLAS and CMS Phase-One Upgrades

LHCb Upgrade

NSF Support for SuperCDMS

MREFC - \$70M TPC and higher (new!)

**Alert!!:** This is only Construction Costs; All Planning Costs must come from program

High-Luminosity ATLAS and CMS detector upgrades

**Input Needed !! Deadline December 8, 2017!!**

NSF 18-013

**Dear Colleague Letter: Request for Information on Mid-scale Research Infrastructure**

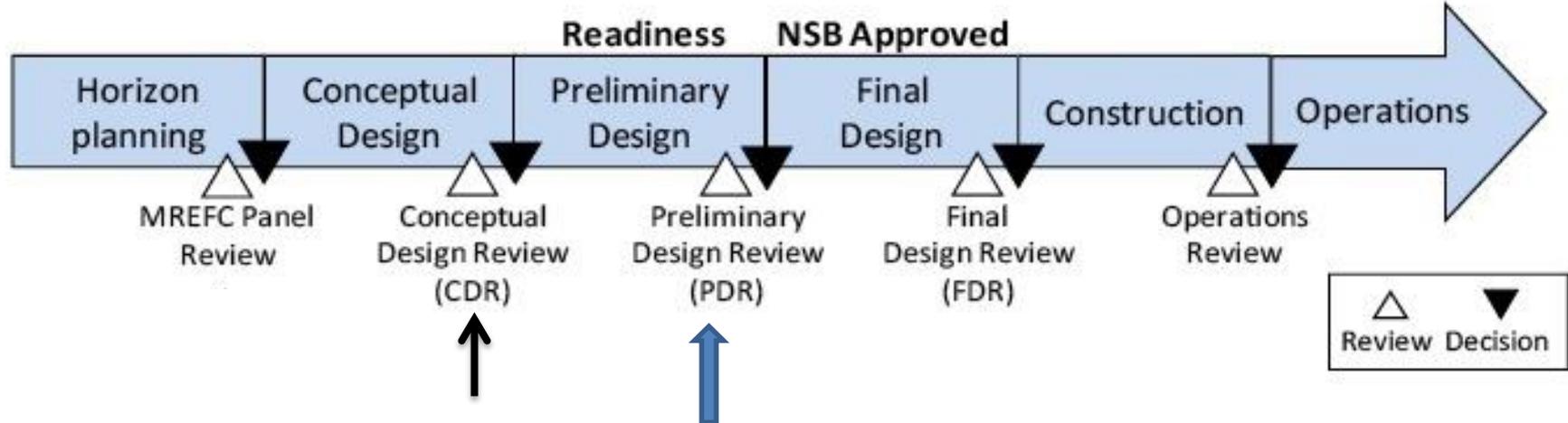
NSF seeks information on existing and future needs for mid-scale research infrastructure projects from the US-based NSF science and engineering community.

Focuses on mid-scale research infrastructure projects with an anticipated NSF contribution of between \$20 million and \$100 million towards construction and/or acquisition



# High-Luminosity LHC Upgrade

Planning for a possible MREFC in support of the high-luminosity upgrades of the ATLAS and CMS detectors at CERN now in Preliminary Design Phase



Coordinating with the DOE in support of the US-CMS and US-ATLAS Teams

PDR sets baseline scope and Total Project Cost for future MREFC request



# Computing

- NSF issued a Software Institute Conceptualization award: [“Conceptualization of an S2I2 Software Institute for High Energy Physics”](#) Award 1558216 (Elmer, Princeton) / 1558233 (Sokoloff, Cincinnati) / 1558233 (Neubauer, UIUC)
- Sponsors community workshops and conceptual work to take advantage of the significant data and computing requirements of the Large Hadron Collider as a science driver for next generation high-performance software and sustainability developments. Working together with the HEP Software Foundation to produce a Community White Paper.
- This effort will inform the future of computing and various software development needs for the HL-LHC era
- **We are partnering with NSF’s Office of Advanced Cyberinfrastructure\***
- Working with ATLAS, CMS, and OSG to minimize disruption to U.S. LHC
- On-going activities and drafts of Community White Paper can be found at: <http://s2i2-hep.org/>

\*Will be important for Harnessing Data (see Big Idea later)



## Major Thrust – Fostering Connections

Focus on Science Question, not Discipline or Subarea

Partner with Others whenever Possible to Promote Science

Partnering within Division – AMO-Nuclear, AMO-Particle, AMO-Gravity

Partnering with other NSF divisions on specific topics –

MPS/AST,CHE,DMR,DMS; BIO/MCB,IOS,DBI; GEO/PLR,AGS;  
ENG/ECCS,CBET; CISE/CCF,OAC

Participation in NSF priority areas jointly with other Directorates/Divisions -  
Understanding the Brain, CIF21, **Big Ideas**

Partnering with DOE in Particle Physics, Nuclear Physics, Plasma Physics

Partnering with NASA in Gravitational Physics and Plasma Physics



## AMO Coupled with PA and Gravitational Physics Very-Low Mass Dark Matter & Gravity at Small Distances

Blewitt and Derevianko, PHY-1506424; “Search of Topological Dark Matter with Atomic Clocks and GPS Constellation”: Analyze clock data from the 30 GPS satellites, which use atomic clocks for everyday navigation. Topological defect dark matter would cause initially synchronized clocks to become desynchronized, causing time discrepancies between spatially separated clocks to exhibit a distinct signature.

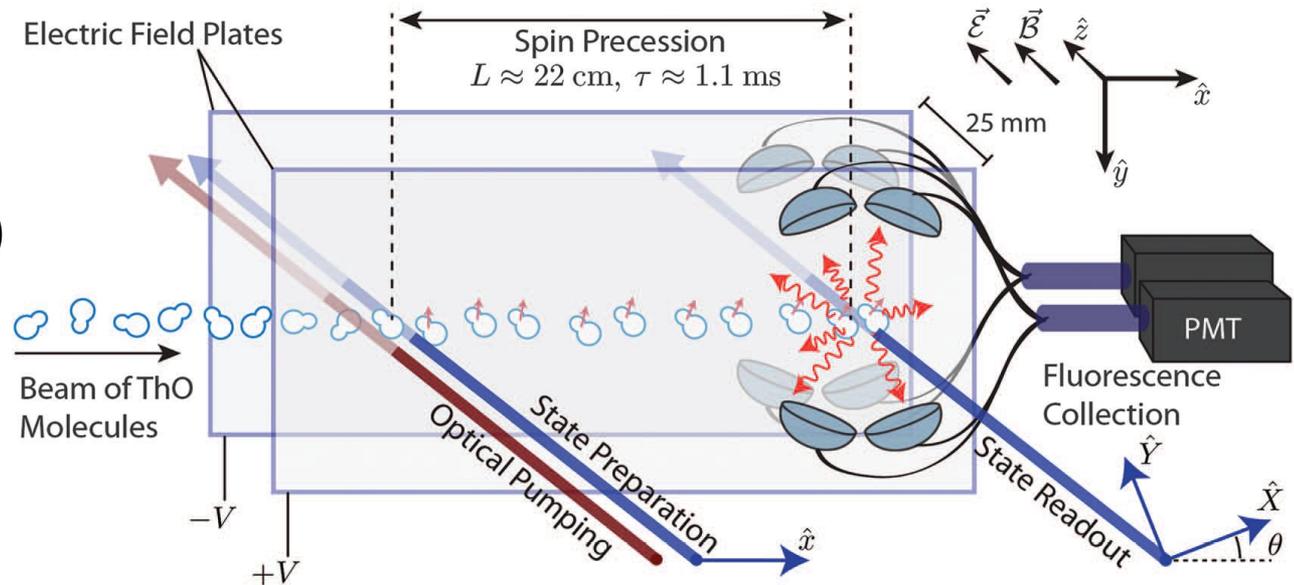


Geraci, PHY-1506431, “Measuring Gravity at the Micron Scale with Laser-Cooled Trapped Microspheres”: Uses levitated spheres of silica to search for hidden forces - sensitivities of a few zeptonewtons. Distinguishes between whether gravity will be much weaker than expected at short range, or stronger.



# ACME: Advanced Cold Molecule Electron Electric Dipole Moment Search

Figure source:  
Science **343**, 269 (2014)



In 2014 the ACME Collaboration placed a limit of  $8.7 \times 10^{-29}$  e-cm on the size of the electron EDM. This measurement is a sensitive probe of new physics at the TeV scale, and complements accelerator-based searches (such as the LHC) for new particles and interactions.

NSF Award 1404146 supports 8 PhD students and 2 postdocs working with 3 PIs to improve this experiments sensitivity to an EDM by a factor of 10.



# Looking Ahead: Ten Big Ideas



**Navigating the New Arctic**



**DATA SCIENCE**  
Fundamental Research - Machine

**Harnessing Data for 21st Century Science and Engineering**



**Work at the Human-Technology Frontier: Shaping the Future**

## RESEARCH IDEAS



**Understanding the Rules of Life: Predicting Phenotype**



**The Quantum Leap: Leading the Next Quantum Revolution**



**Windows on the Universe: The Era of Multi-messenger Astrophysics**

## PROCESS IDEAS



**Growing Convergent Research at NSF**



**NSF-Includes: Enhancing Science and Engineering through Diversity**



**Mid-scale Research Infrastructure**



**NSF 2050: Seeding Innovation**



## 10 Big Ideas for Future NSF Investments

- Bold questions that will drive NSF's long-term research agenda
- Catalyze investment in fundamental research
- Collaborations with industry, private foundations, other agencies, universities
- Solve pressing problems and lead to new discoveries



# Windows on the Universe

The goal of “Windows on the Universe” is to bring electromagnetic waves, high-energy particles, and gravitational waves together to study the universe and probe events in real time in a way that was previously impossible.



Credit: LIGO Laboratory



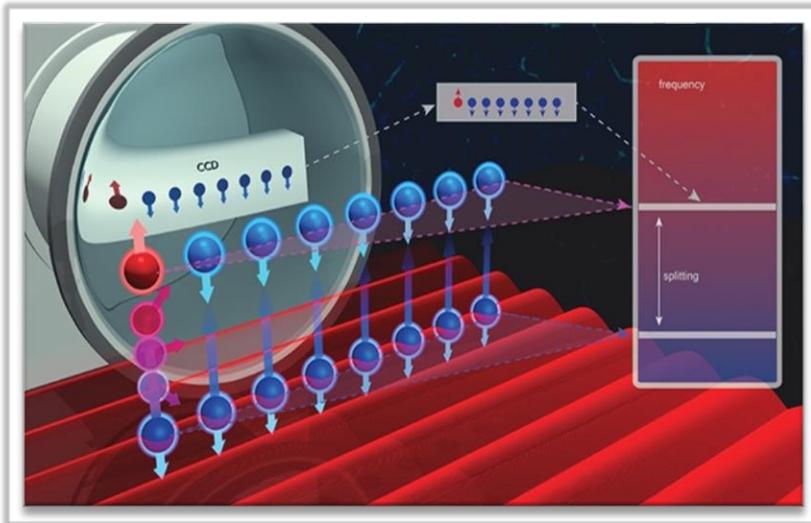
Credit: IceCube



Credit: AURA



# Quantum Leap



Trapped ion computation  
(JQI – University of Maryland)

## Today:

- lasers, atomic clocks, GPS, semiconductors, storage media

## Tomorrow:

- Ultra-secure communication
- Ultra-precise sensing, measurement
- Quantum simulators
- Computing beyond the scale of supercomputing



## Builds on 20-Year NSF Investment in Quantum Information Science

“Advancing Quantum Information Science: National Challenges  
and Opportunities”, NSTC, July 2016

Quantum Leap is an NSF-Wide Activity Involving Multiple Directorates  
that extends QIS into new regimes of materials science and engineering

NSF/DOE Quantum Science Summer School, JHU, 5-16 June 2017  
Three more to follow: Cornell (2018); Penn State (2019); UC (2020)

Triplets: Quantum Information Science and Engineering Network, PI Awschalom

NSF 17-053 Funding Opportunity – A “Quantum Leap” Demonstration of  
Topological Quantum Computing

Ideas Lab: Practical Fully-Connected Quantum Computer Challenge (PFCQC)  
Ideas Lab in August; Proposals due November 30



# Harnessing Data

Domain Science  $\leftrightarrow$  Computation and Data Challenges

Connecting domain scientists with computer scientists

Essential for Handling Large Data Sets from LIGO, LHC, LSST

Efforts closely coordinated with OAC in CISE Directorate