

# REPORT TO THE NUCLEAR SCIENCE ADVISORY COMMITTEE

## *Neutrinoless Double Beta Decay*

<http://science.energy.gov/np/nsac/reports/>

### *Briefing to HEPAP*

**Robert D. McKeown**

# Outline

- Charge and Membership
- Science Overview and Update
- Current and Proposed Projects
- R&D Plans
- Summary

# 2015 Charge

...the NSAC Subcommittee on Neutrinoless Double Beta Decay is requested, in the context of ongoing and planned US efforts as well as international competitiveness, to consider the following:

- Assess the status of ongoing R&D for NLDBD candidate technology demonstrations for a possible future ton-scale NLDBD experiment.
- For each candidate technology demonstration, identify the major remaining R&D tasks needed ONLY to demonstrate downselect criteria, including the sensitivity goals, outlined in the NSAC report of May 2014. R&D needs for candidate technology demonstrations should be sufficiently documented beyond assertion to allow critical examination by the panel and future assessments.
- Identify the time durations needed to accomplish these activities and the corresponding estimated resources, as reported by the candidate technology demonstration groups.

# Further Clarification

Guided by the R&D assessments provided by this NSAC subcommittee and within funding availability, NSF and DOE plan to move forward in a coordinated, unified approach to address these R&D needs, similar to the process used in the recent joint effort on the second generation dark matter experiments. That process included independent calls for proposals with coordinated requirements, and a joint review. The role of this NSAC subcommittee is to assess the proposed R&D plans for each technology, but not to review the relative value of the plans for one technique vs. another.

*(Text from our report, following discussion with NSF/DOE)*

# Subcommittee Membership

R. McKeown (Chair)

F. Calaprice

V. Cirigliano

P. Cushman

D. Geesaman (ex-officio)

G. Greene

J. Hardy

D. Hertzog

M. Kamionkowski

K. Langanke

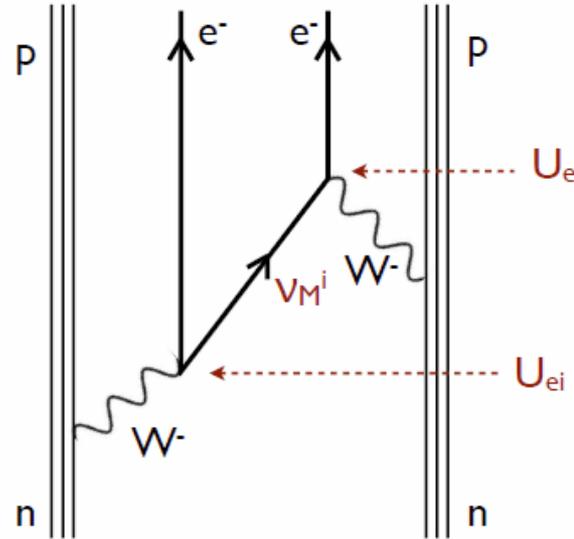
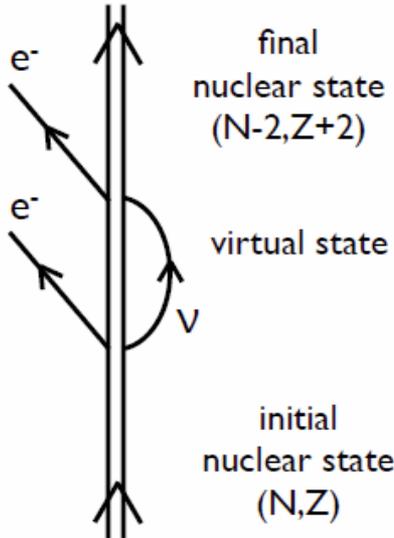
K. Scholberg

H. Sobel

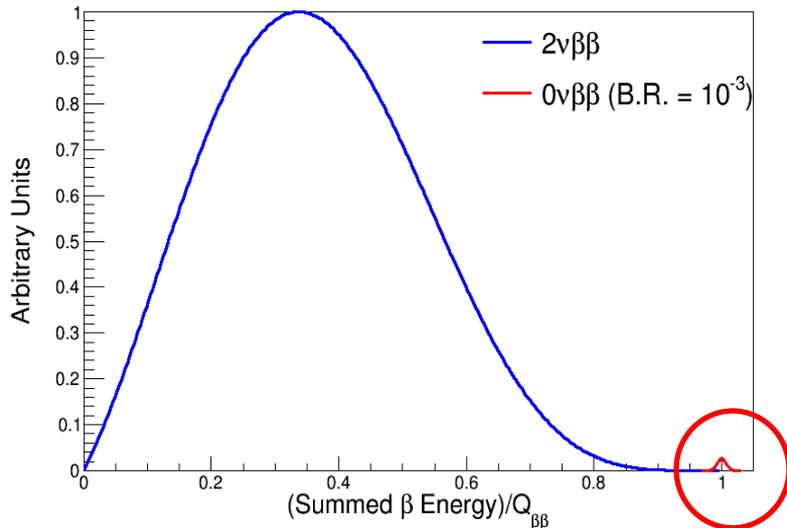
S. Vigdor

- Representation from NP/HEP-exp, NP/Astro-theory
- One meeting in DC, August 17-19 with presentations from ongoing projects
- Report to NSAC October 15, 2015 – slightly revised final report dated November 18

# $0\nu\beta\beta$ decay



- Majorana  $\nu$
- Flip helicity:
  - RH coupling
  - $m \neq 0$
- $\Delta L = 2$



## Experimental Issues

- Good energy resolution
- Low background
- ➔ R&D

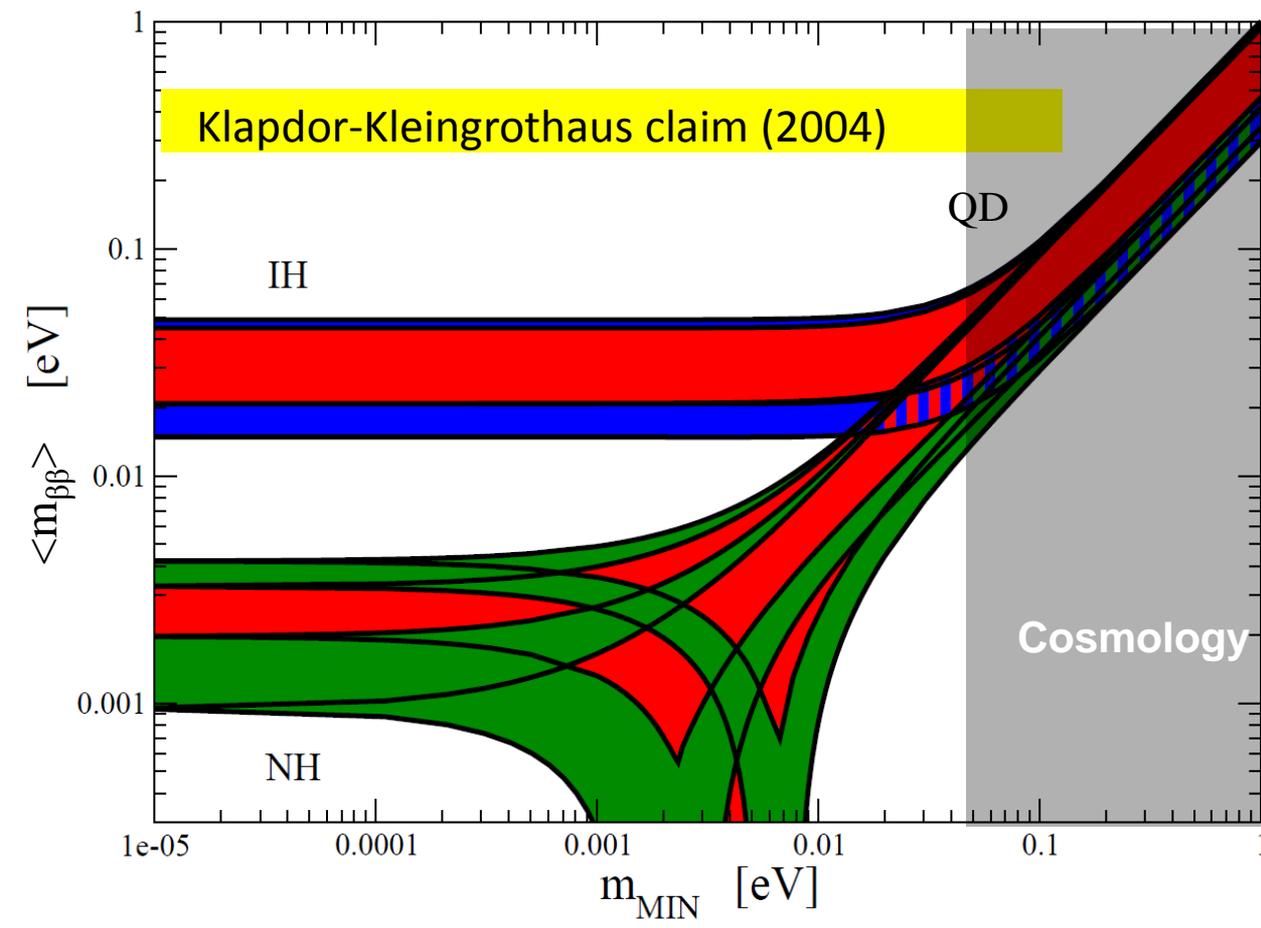
# NLDBD and Neutrino Mass

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu} \cdot |M^{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$$

Phase space

Nuclear Matrix Element

$$\langle m_{\beta\beta} \rangle^2 = \left| \sum_i U_{ei}^2 m_{\nu_i} \right|^2$$

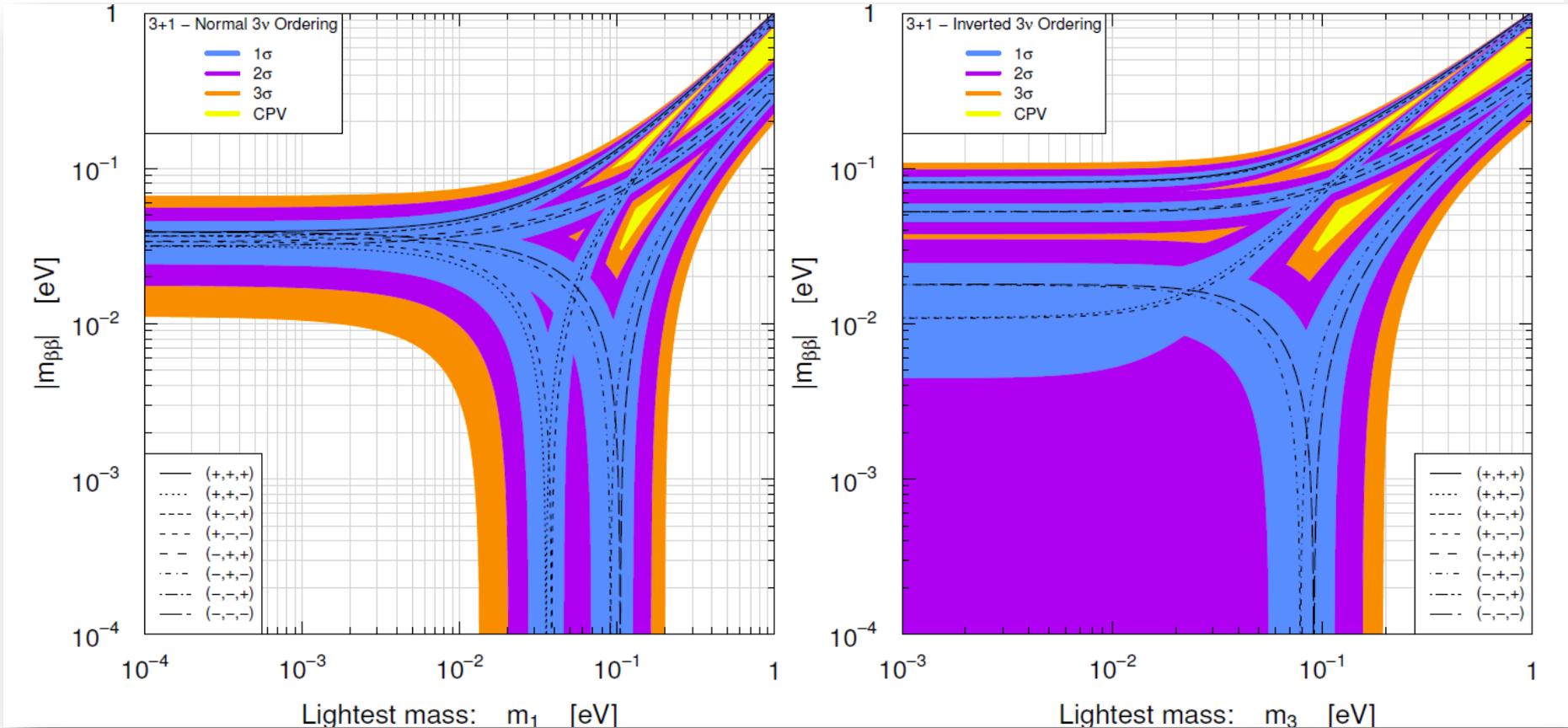


Note: colored bands  
Indicate allowed  
variation of  $U_{ei}$  due to  
unknown Majorana  
phases and uncertainty  
in mixing angles

- $\langle m_{\beta\beta} \rangle^2 = \left| \sum_i U_{ei}^2 m_{\nu_i} \right|^2$
- $m_{\text{MIN}} = \text{lightest } m_{\nu_i}$

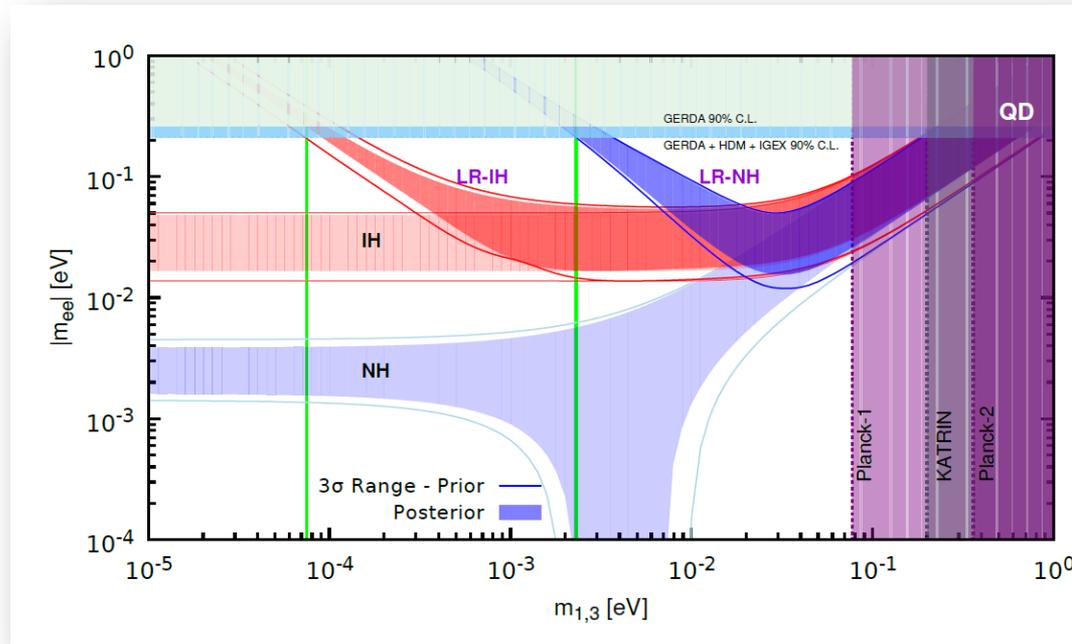
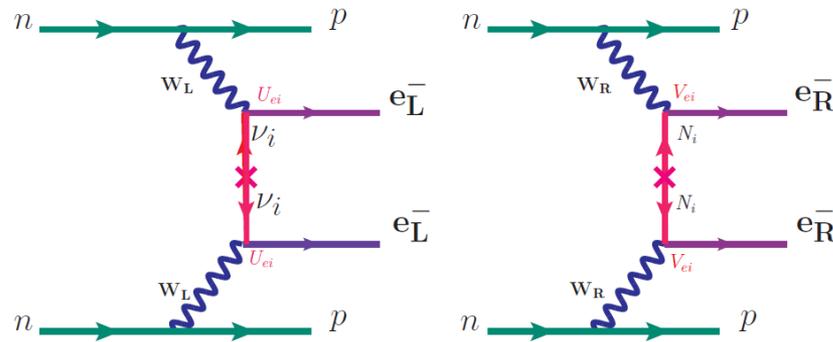
# Sterile Neutrinos and NLDBD

$$|m_{\beta\beta}| = \left| |U_{e1}|^2 m_1 + |U_{e2}|^2 e^{i\alpha_2} m_2 + |U_{e3}|^2 e^{i\alpha_3} m_3 + |U_{e4}|^2 e^{i\alpha_4} m_4 \right|$$



arXiv:1507.08204

# New Physics and LHC

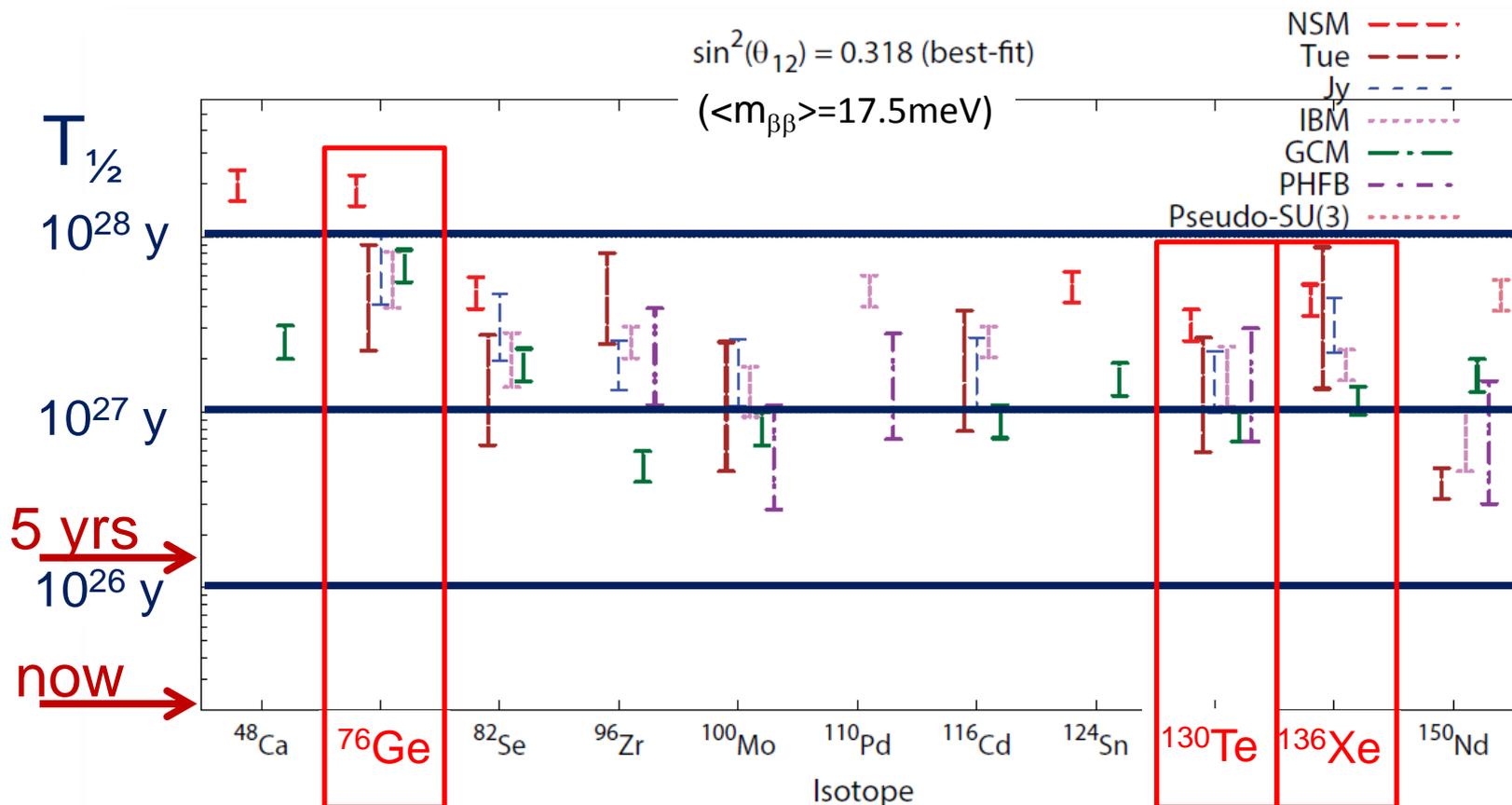


arXiv:1508.07286

# Science Assessment

“...it is important to remember that NLDBD has a unique role in potentially addressing the issue of Dirac vs. Majorana nature of neutrinos. The Subcommittee remains convinced that the scientific case for pursuing NLDBD experiments at the ton-scale is very compelling.”

# Inverted Hierarchy Coverage



$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu} \cdot |M^{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$$

Figure source: A. Dueck, W. Rodejohann, and K. Zuber, Phys. Rev. D83 (2011) 113010.

# Nuclear Theory Developments

- Application of modern techniques to  $0\nu\beta\beta$  and  $2\nu\beta\beta$ 
  - Ab initio methods
    - 1.) Light nuclei to test  $g_A$  quenching for  $0\nu\beta\beta$
    - 2.) Develop better effective interactions for heavier nuclei
  - Better approximations for heavy nuclei
    - 1.) Larger model spaces
    - 2.) Density Functional Theory
    - 3.) Interacting Boson Model
- Larger and broader group of nuclear theorists interested in working on this problem
- Dedicated “Topical Collaboration” selected for funding by DOE-NP

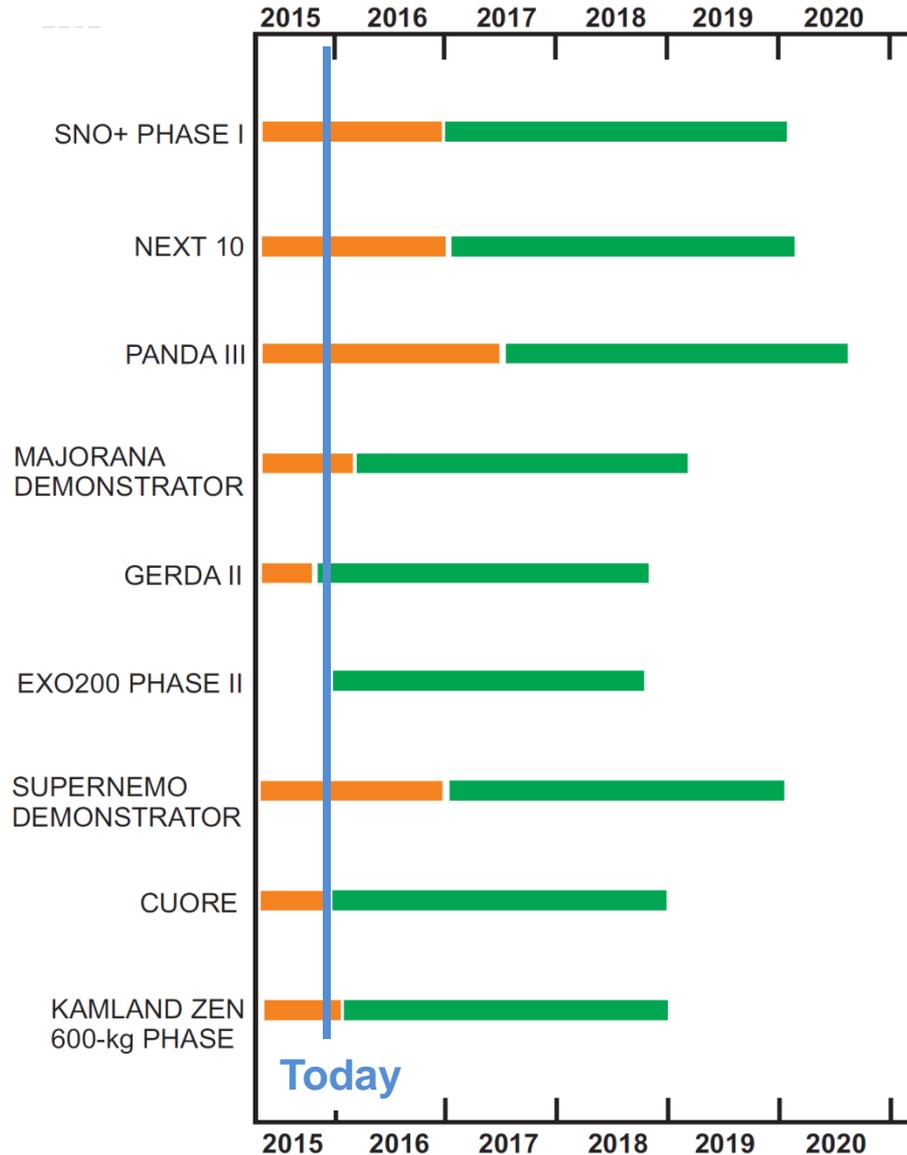
# Current Projects

Project	Isotope	Isotope Mass (kg fiducial)	Currently Achieved ( $10^{26}$ yr)	Location	US Funding Level
CUORE	$^{130}\text{Te}$	206	>0.028	Gran Sasso	Major
MAJORANA	$^{76}\text{Ge}$	24.8		SURF	Major
GERDA	$^{76}\text{Ge}$	31	>0.21	Gran Sasso	None
EXO-200	$^{136}\text{Xe}$	79	>0.11	WIPP	Major
NEXT	$^{136}\text{Xe}$	10→100		Canfranc	Minor
SuperNEMO	$^{82}\text{Se}+$	7	>0.001	Frejus	Minor
KamLAND-Zen	$^{136}\text{Xe}$	434	>0.19	Kamioka	Minor
SNO+	$^{130}\text{Te}$	160		SNOLAB	Minor
PANDAX-III	$^{136}\text{Xe}$	200		Jinping	Minor

## Primary goals:

- Demonstrate background reduction for next generation experiment
- Extend sensitivity to  $T_{1/2} \sim 10^{26}$  years.

# Updated Timeline



Construction  
Operation  
(not time until downselect)

# Subcommittee Observation

One can see that there is about 1 more year of construction and assembly before all the projects are in an operational phase taking data. Therefore, over the next 2-3 years one can expect to have valuable information based on real data combined with results from additional R&D for these different techniques. At that point, one would expect that an assessment of the relative merits would be more reliable than at the present time.

# Overall R&D Assessment

In general, the suite of mid-scale experiments and demonstration projects are making good progress in setting new  $0\nu\beta\beta$  limits and in testing out techniques that can be extrapolated to ton-scale installations. During the next 1-2 years many techniques will be acquiring data and producing a body of information that will inform the future plans. However, it is already clear that additional R&D issues must be resolved in preparation for a future downselect decision. **Therefore the subcommittee strongly recommends that R&D efforts aimed at solving specific technical issues relevant to the downselect decision be supported.**

# Proposed U.S. R&D

<b>Germanium</b>	Higher radiopurity connectors	Radiopure fabrication methods	
<b>CUPID</b>	Particle ID	Component radiopurity	Study Cosmogenics with CUORE
<b>NEXT</b>	Study NEXT10, DEMO, DEMO+	Fluorescence Ba detection	
<b>PANDA X III</b>	HV cage and radiopure pressure vessel	Topmetal readout	
<b>SNO+ Phase II</b>	Run SNO+	Load Te to 3%	Increase Light yield
<b>nEXO</b>	HV @ >50kV	Hi QE radiopure photodetectors (SiPM)	Electronics
<b>SuperNEMO</b>	Foil radiopurity	Scintillator/PMT improvement	Tracker improvement

Total estimated resources ~\$11M

# Additional Statement on R&D

Other technical issues have more open-ended R&D requirements to address. In these cases the allocation of resources will be more difficult to assess. In any case, the longer term future of NLDBD will require continued R&D effort. **The subcommittee strongly urges continuation of longer term R&D necessary for the future development of the subject in addition to the support of shorter term R&D aimed at a near future downselect.**

# Common R&D Topics

It was noted by the subcommittee that there are several common R&D topics that would benefit several different techniques. It seems in these cases that a coordinated approach could be a more efficient use of resources. **The subcommittee suggests that the funding agencies consider an approach that would encourage several groups to work together on these common goals.**

# International Aspects

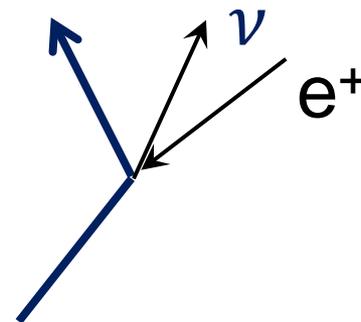
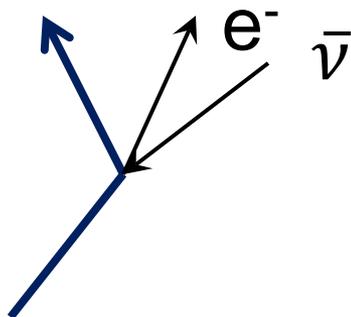
We continue to advocate that the US should plan for a leadership role in (at least) one experiment, while perhaps maintaining options to participate in one or more internationally led projects. This will require timely and astute assessment of both the technological opportunities as well as the inherent strengths of research groups in various countries. At this point, the best one can say is that it is advisable to maintain a nimble posture, with an eye towards a timely decision in the near future (perhaps as short as 2 years).

# Acknowledgements

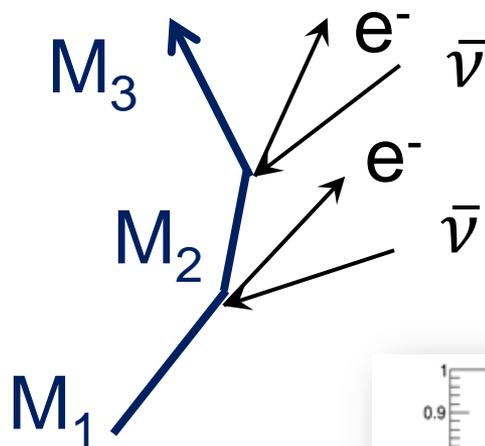
- Thanks to the collaborations for providing valuable material that was essential to the Subcommittee in its work.
- Thanks to SURA for hosting the open meeting in August.
- Thanks to Brenda May (DOE-NP) and Pat Stroop (JLab) for logistical assistance.
- Thanks to Mary Beth Stewart (JLab), and others, for assistance in preparing our report.
- Thanks to Subcommittee members for diligent efforts on a very challenging time frame.

# Backups

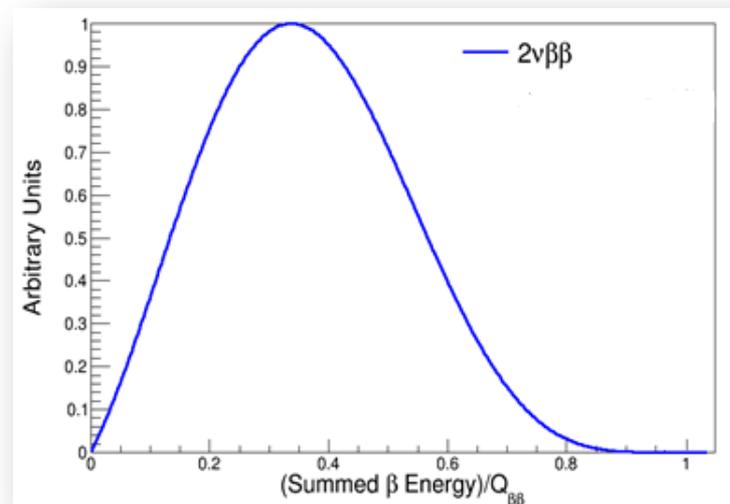
- $\beta^\pm$  decay



- $2\nu\beta\beta$  decay

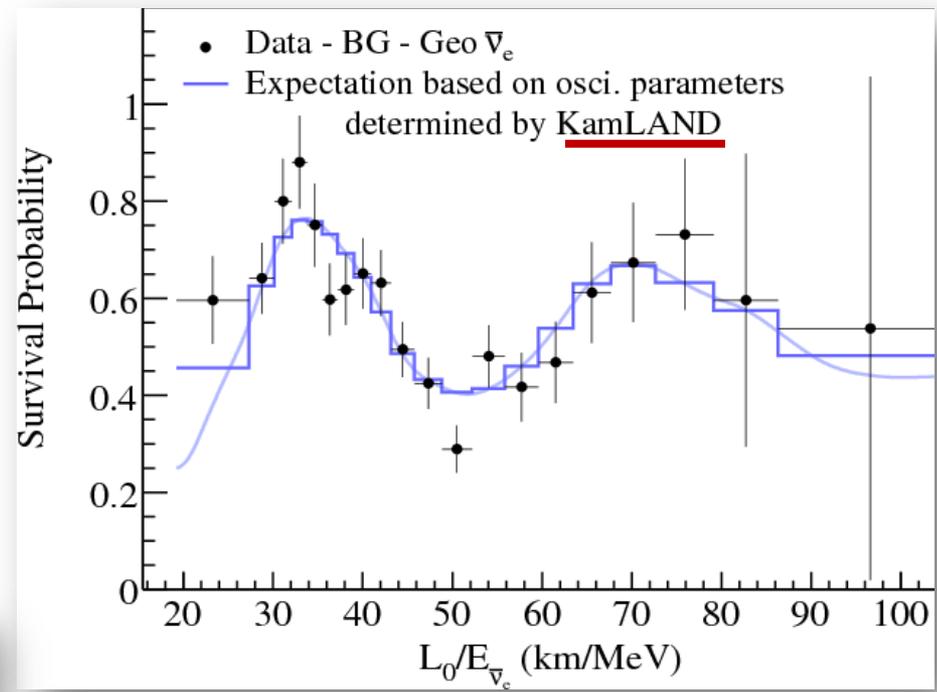
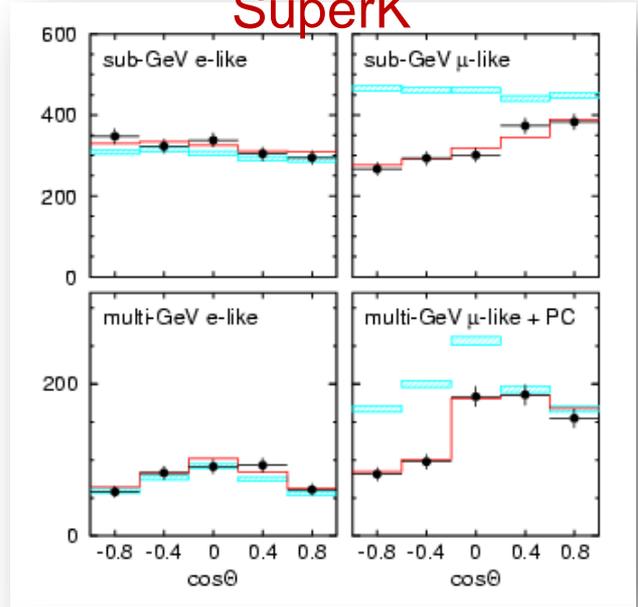


- $M_2 > M_1 > M_3$
- 2<sup>nd</sup> order weak process

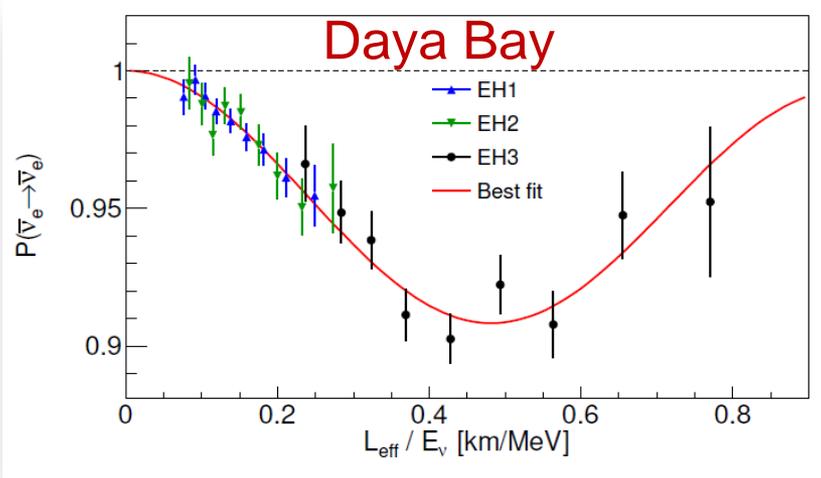


# Neutrino Oscillations ( $m_\nu \neq 0$ )

SuperK



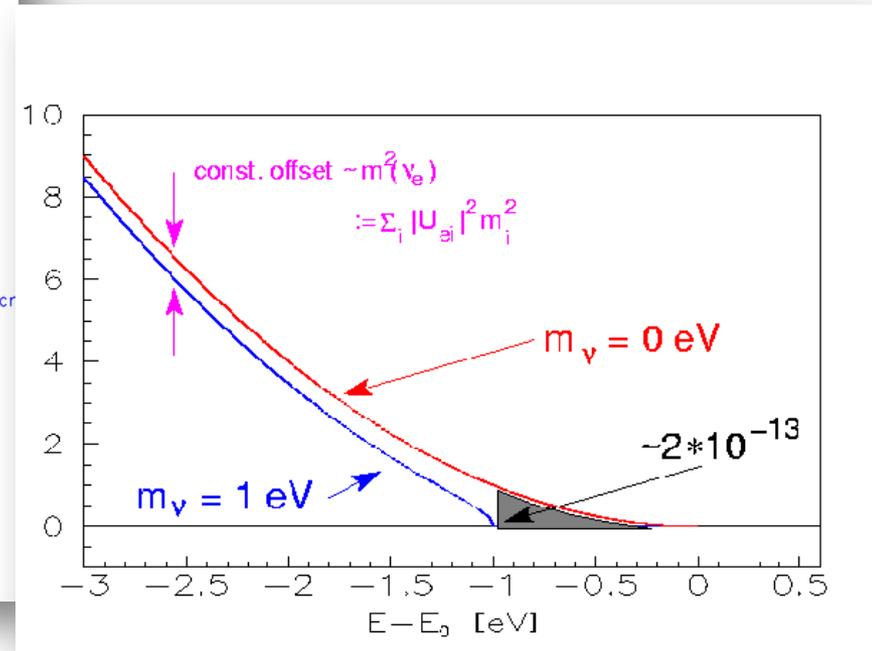
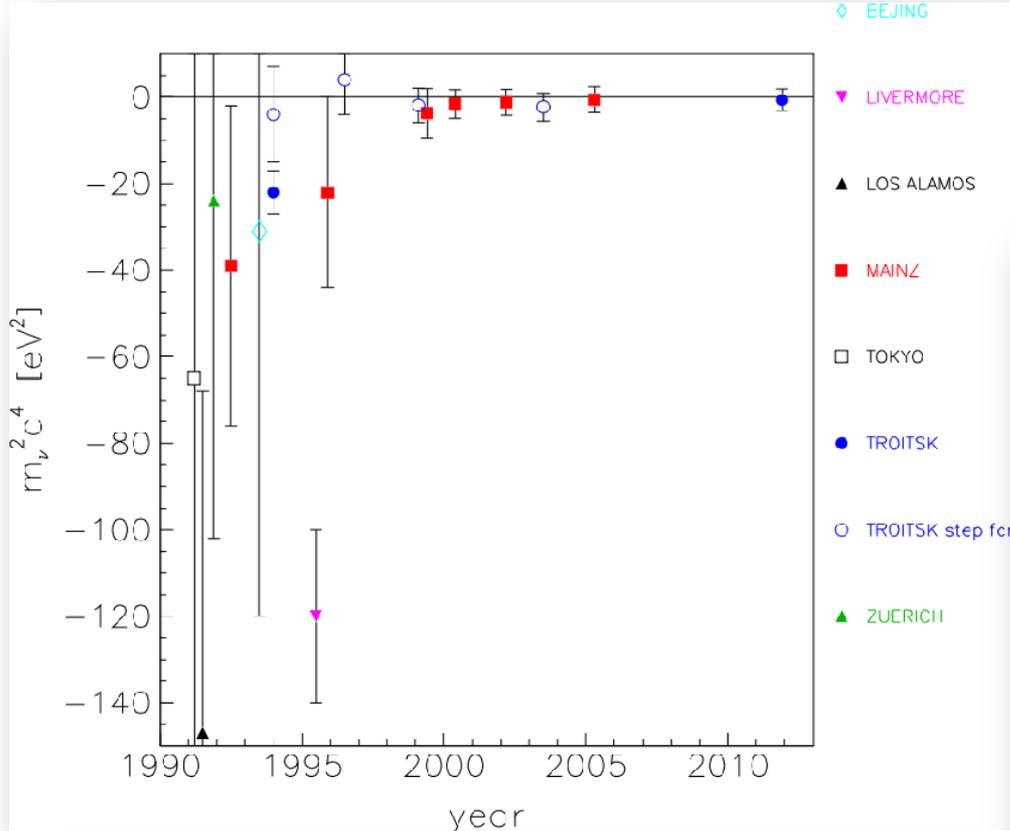
Daya Bay



Measure:

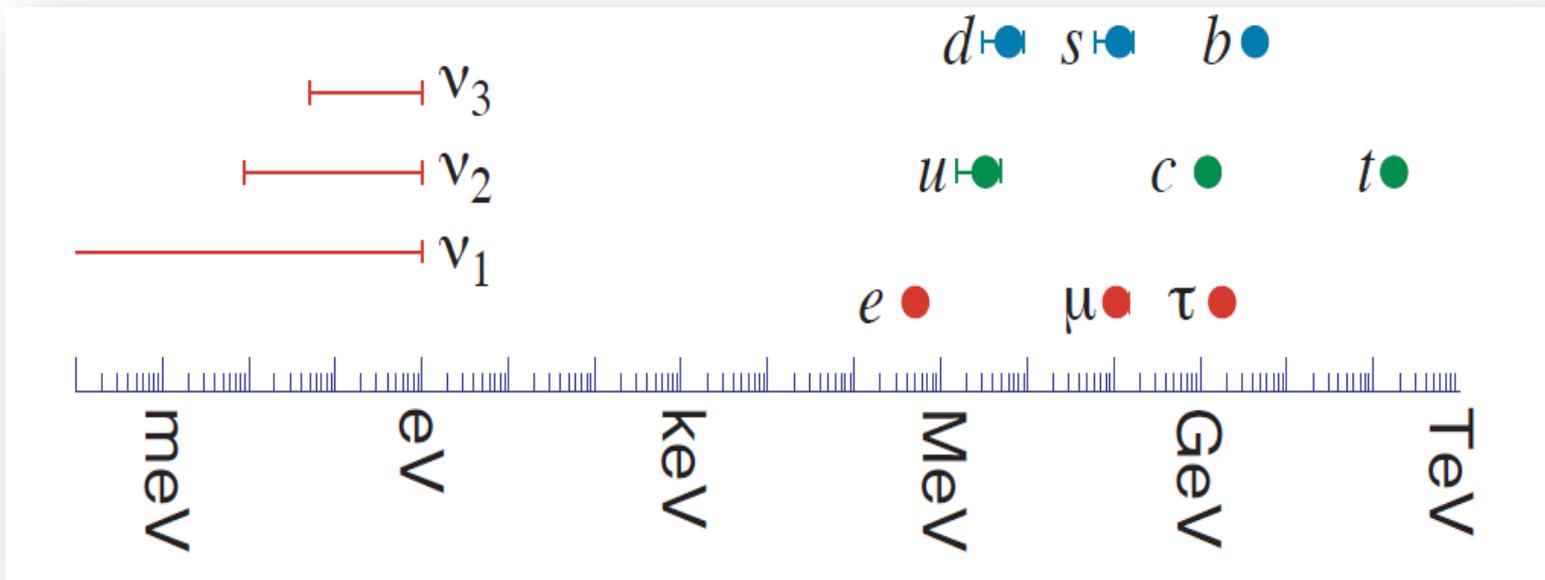
- $|\Delta m_{ij}^2|$
- $\theta_{ij}$

# Absolute Neutrino Mass Limits



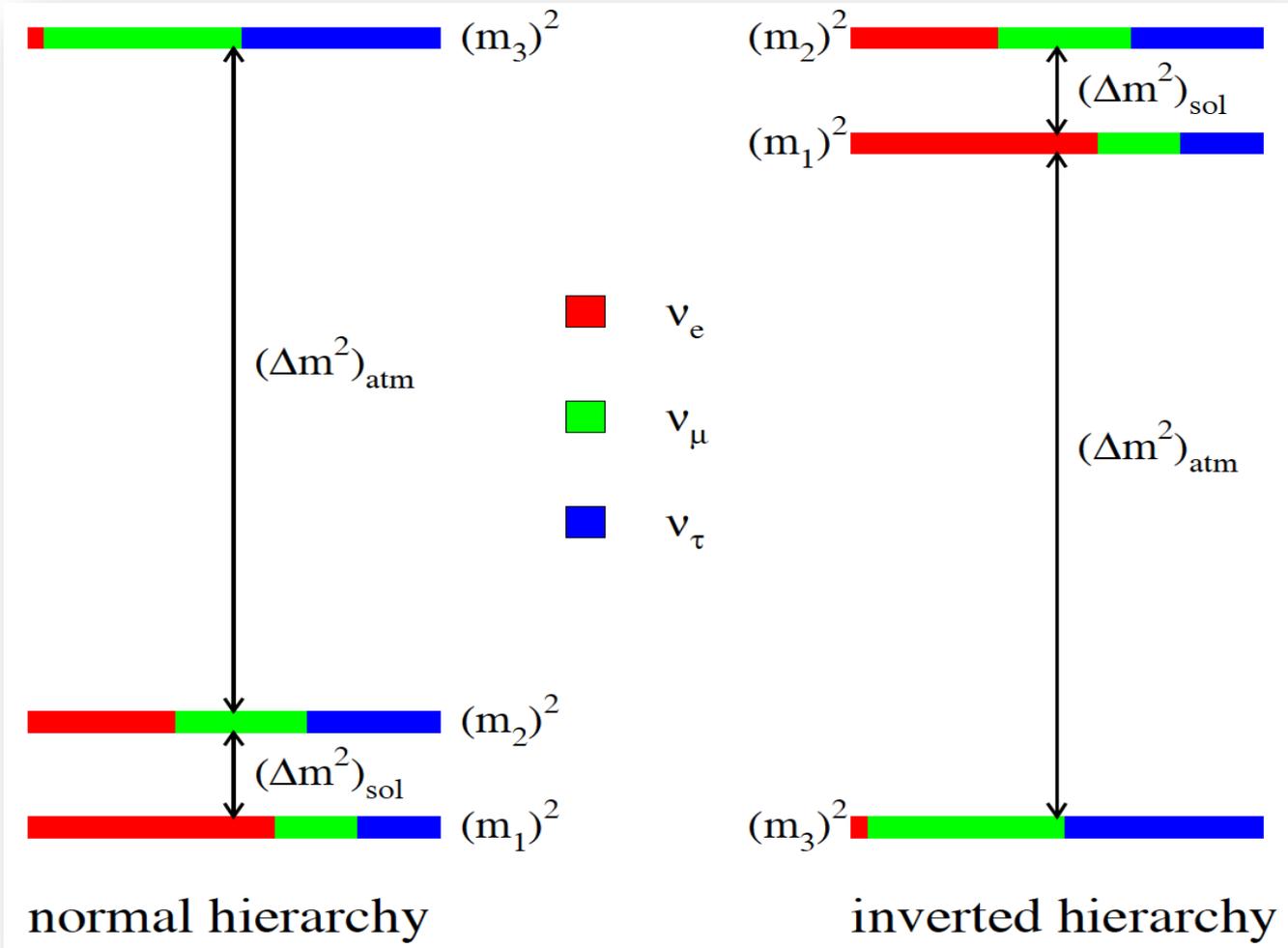
- Present limit from tritium decay:  $< 2 \text{ eV}$
- Cosmology:  $\sum m_i < 0.23 \text{ eV}$  (95% CL)

# Masses of Matter particles



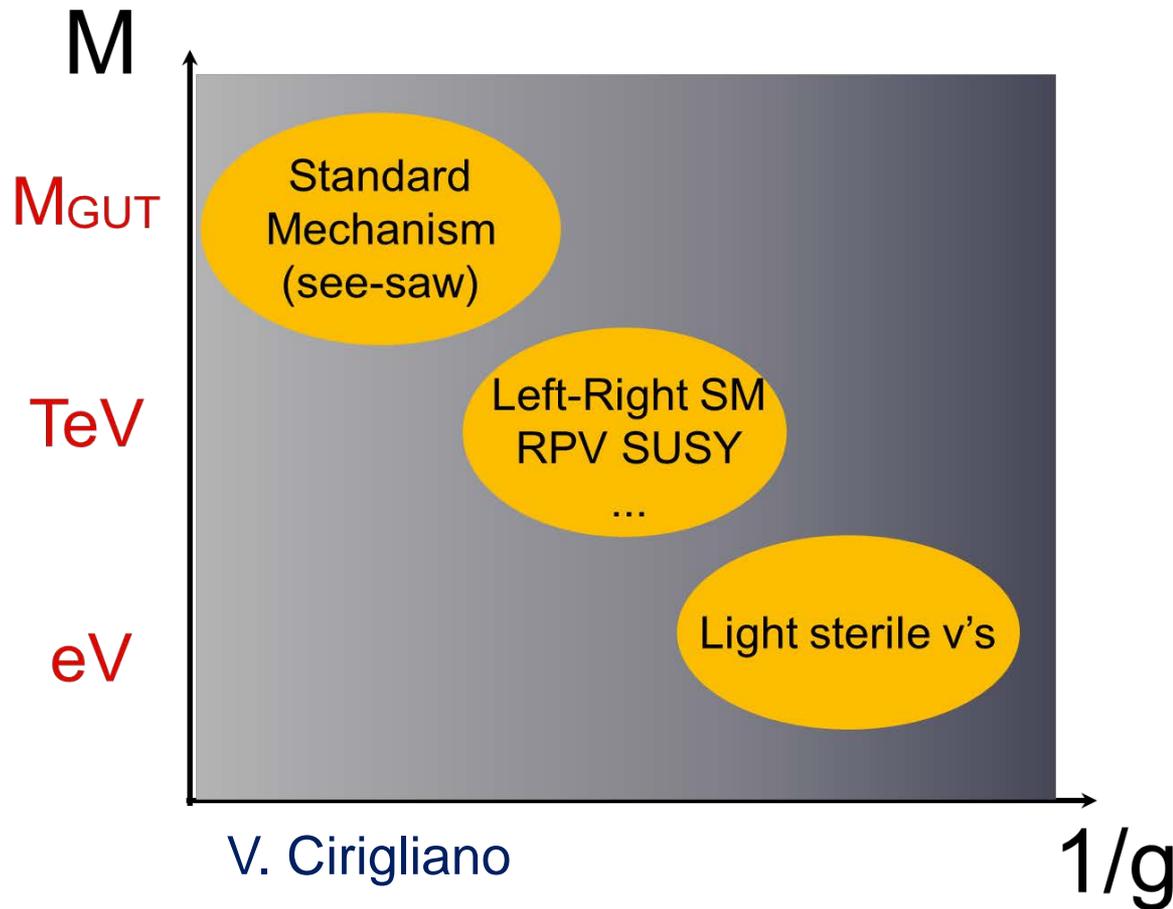
- Higgs mechanism not likely responsible for neutrino masses
- “See-saw” is most common alternative
  - Majorana neutrinos!
  - Leptogenesis

# Mass Hierarchy



# Lepton Number Violation and $0\nu\beta\beta$

- Ton-scale  $0\nu\beta\beta$  probes LNV from a variety of mechanisms and scales of masses ( $M$ ) and couplings ( $g$ )



# Simple Background Estimate

$$\text{NLDBD Rate} = N \times \ln(2) / T_{1/2} \text{ (assume } T_{1/2} \approx 10^{28} \text{ yr)}$$

$$\text{For 1 Tonne, } N = 10^6 \text{g} \times 6 \times 10^{23} / \text{MW}$$

(MW = 67, 130, 136 → use MW ≈ 100)

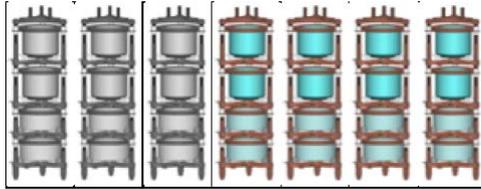
$$\text{So } N \approx 6 \times 10^{27}$$

$$\text{NLDBD Rate} = 0.4 / \text{Tonne/yr}$$

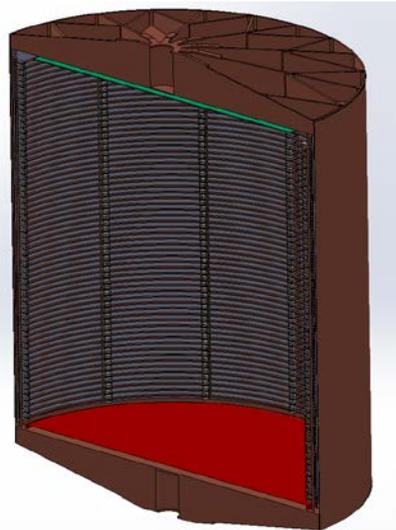
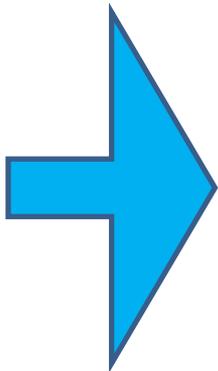
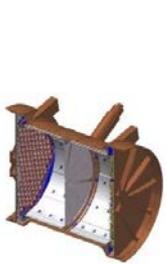
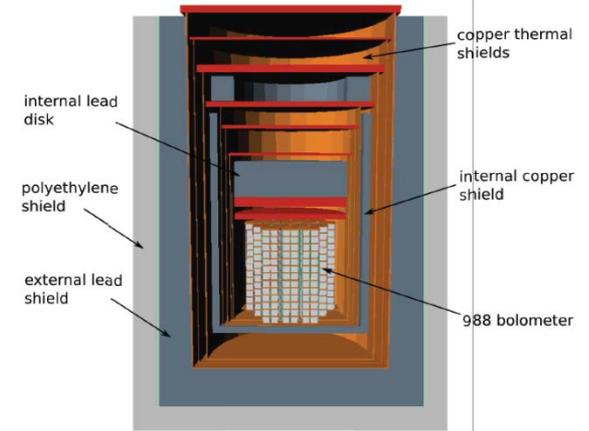
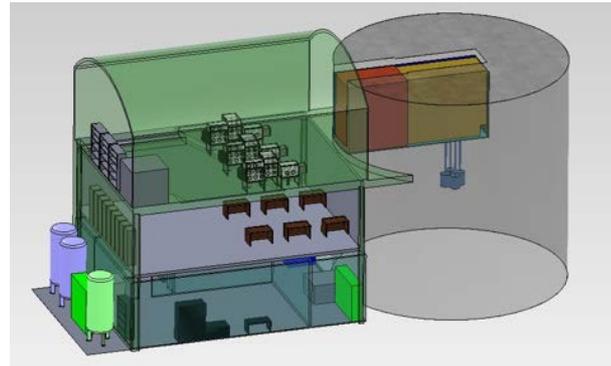
$$\text{Background free} \rightarrow \text{Background} < 0.1 / \text{Tonne/yr/ROI}$$

# Next Generation Approaches

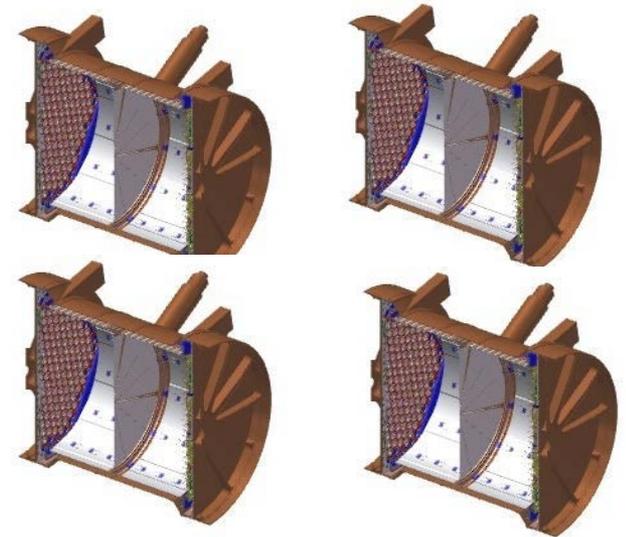
The issue is to scale up to  $\geq 1$  Tonne with low background.



X ~100 →



OR



# Technology Assessment

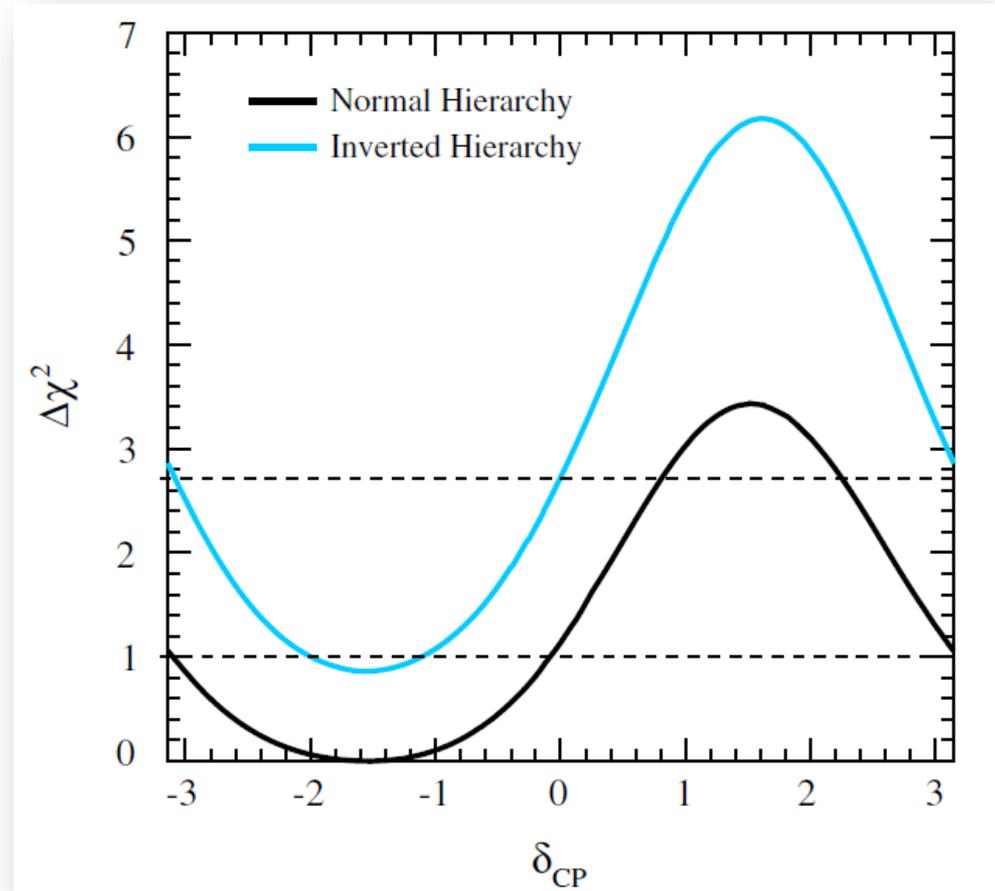
For each project/technique, our report has the structure:

- Status
- Plans for R&D (summary)
- Technical Issue 1
  - Proposed R&D
  - Comments
- Technical Issue 2
  - Proposed R&D
  - Comments
- Other Technical issues (generally beyond downselect)

Note: Sent to collaborations for fact check

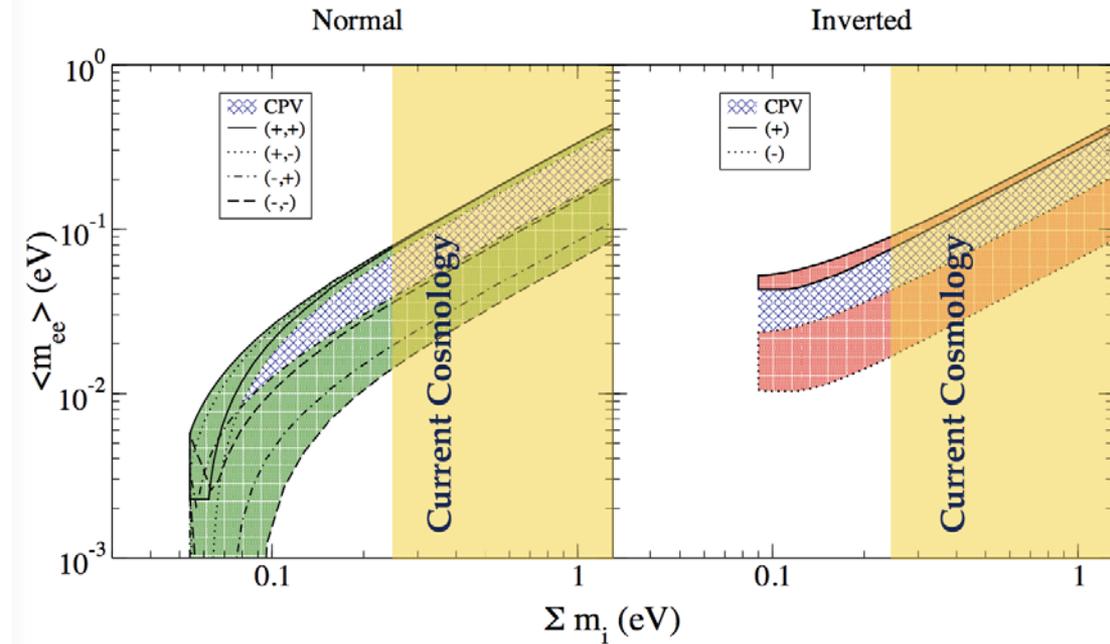
# Neutrino Oscillation Experiments

- T2K reports the result shown (combined fit with reactor expts)
  - favors  $\delta_{CP} \sim -\pi/2$
  - slightly favors NH
- First results from NovA at Fermilab are consistent
- Both keep running...
- PINGU, JUNO, RENO50 all aim for mass hierarchy within a decade



Phys.Rev. D91 (2015) 7, 072010

# Cosmological Limits

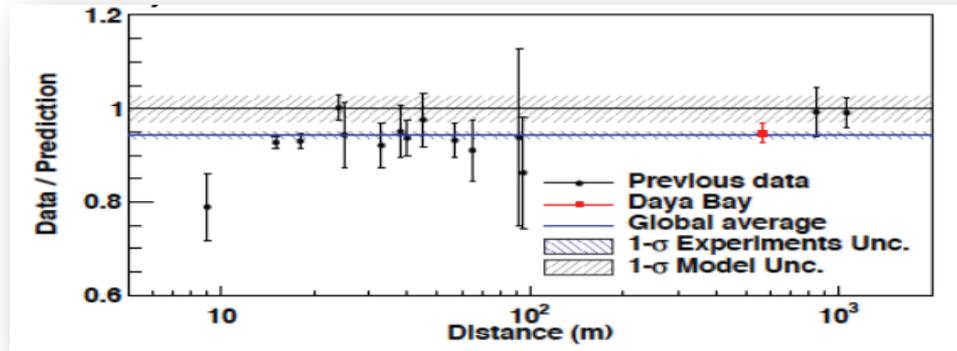


- Within the  $\Lambda$ CDM model, cosmology probes the neutrino freestreaming scale (neutrino hot dark matter component), which depends on  $\Sigma \equiv \Sigma_i m_i$  and the relic neutrino energy spectra
- Current combined bound:  
 $\Sigma < 230$  meV
- Projected bounds (<10 years):  
 $\Sigma < 100$  meV (could indicate ordering)

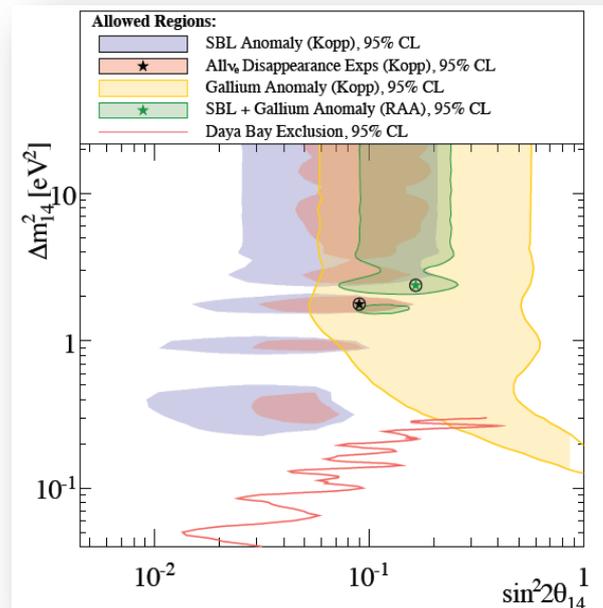
Presentation to Subcommittee by K. Azerbaijan (C Irvine)

# Sterile Neutrinos

- Reactor Anomaly:



- Fits to sterile  $\nu$ 's:



Many experiments being mounted and proposed to confirm/refute interpretation

# Projects Data Collection

- Requested documentation on
  - Current status and plans
  - R&D required for downselect
- Scheduled 7 presentations for August 17-19 open meeting at SURA HQ in DC

Note: all submissions and slides are kept private for Subcommittee use

# Methods

- $^{136}\text{Xe}$  TPCs (liquid, gas)
- $^{76}\text{Ge}$  Crystals
- Bolometers with particle ID enhancements
- Doped Liquid Scintillators ( $^{136}\text{Xe}$ , Te)
- Foils with tracking chambers ( $^{82}\text{Se}$  + )

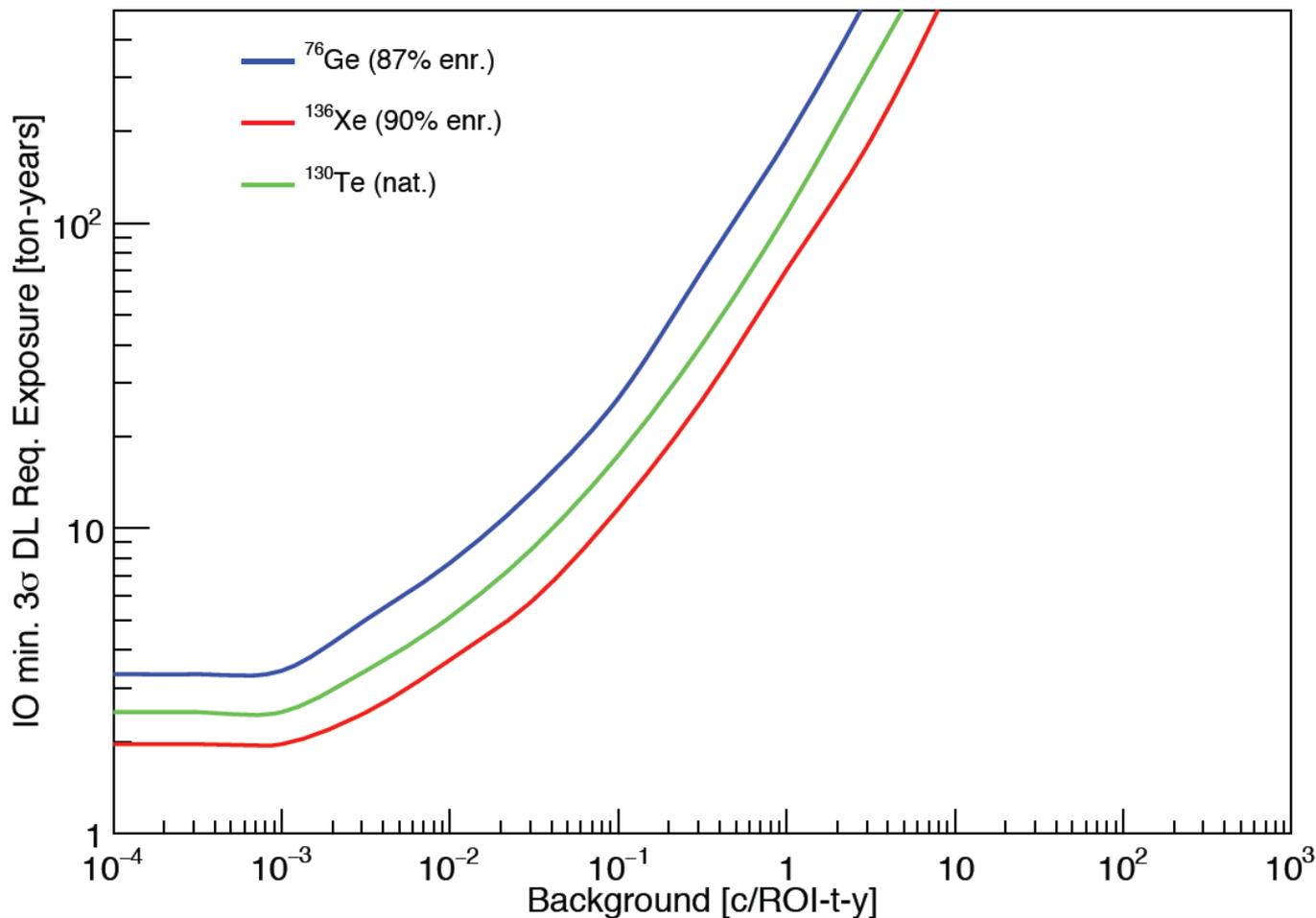
# Major Issue: Background

- For “background-free” experiment, lifetime sensitivity goes as  $T_{1/2} \sim M \cdot t_{\text{run}}$   
( $M$ = isotope mass)  
→ factor of 50 in  $T_{1/2}$  needs factor of 50 in  $M$  (for constant  $t_{\text{run}}$ )
- For experiment with background, as  $T_{1/2} \sim (M \cdot t_{\text{run}})^{1/2}$   
→ factor of 50 in  $T_{1/2}$  needs factor of 2500 in  $M$  (for constant  $t_{\text{run}}$ )
- Background reduction is the key to a successful program
  - deep underground
  - radiopurity
  - better E resolution
  - better event characterization

→ R&D will be crucial

# Required $3\sigma$ Exposure vs. Background

J. Detwiler



“Required” exposure assuming minimum IO  $m_{\beta\beta}=18.3$  meV, taken from using the PDG2013 central values of the oscillation parameters, and the most pessimistic NME for the corresponding isotope among QRPA, SM, IBM, PHFB, and EDF

(IO = Inverted Ordering)