



# Dark Energy Survey

HEPAP Summer Meeting  
28 August 2012

DARK ENERGY  
SURVEY



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# Dark Energy Survey Concept

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- **Survey concept:**
  - Perform a 5000-square-degree, 5-band,  $\sim 24^{\text{th}}$  magnitude survey of the southern sky
- **New instrument:**
  - Replaced this prime focus cage with a 2.2 degree FOV, 570 Megapixel red-sensitive CCD imager and optics
- **Time scale:**
  - Preparation and reviews 2003-2008
  - Instrument construction 2008-2011
  - Delivery to CTIO 2011-2012
  - Installation Feb.- Aug. 2012
  - Commissioning Sept. 2012
- **525-night survey:**
  - 105 nights per year (Sept.- February)

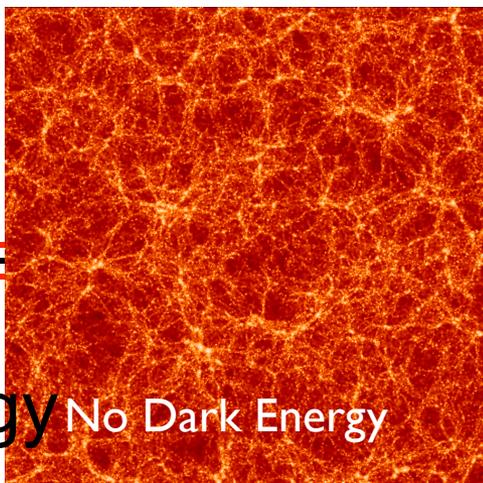


Use the Blanco  
4-m Telescope  
at the Cerro Tololo  
Interamerican  
Observatory (CTIO)

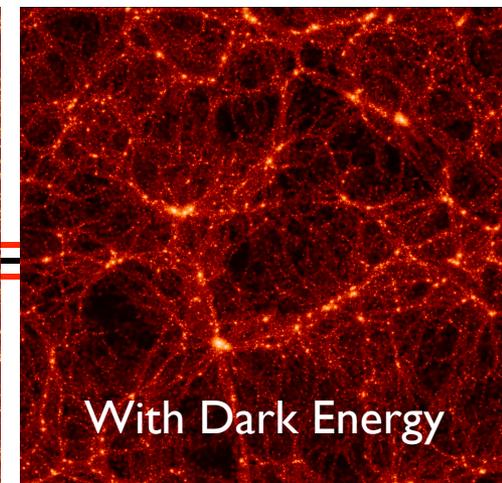


# DES Science Summary

DARK ENERGY SURVEY



No Dark Energy



With Dark Energy

## Four Probes of Dark Energy

### • Galaxy Clusters

- ~100,000 clusters to  $z = 1$  and beyond
- Synergy with South Pole Telescope
- Sensitive to growth of structure and geometry

### • Weak Lensing

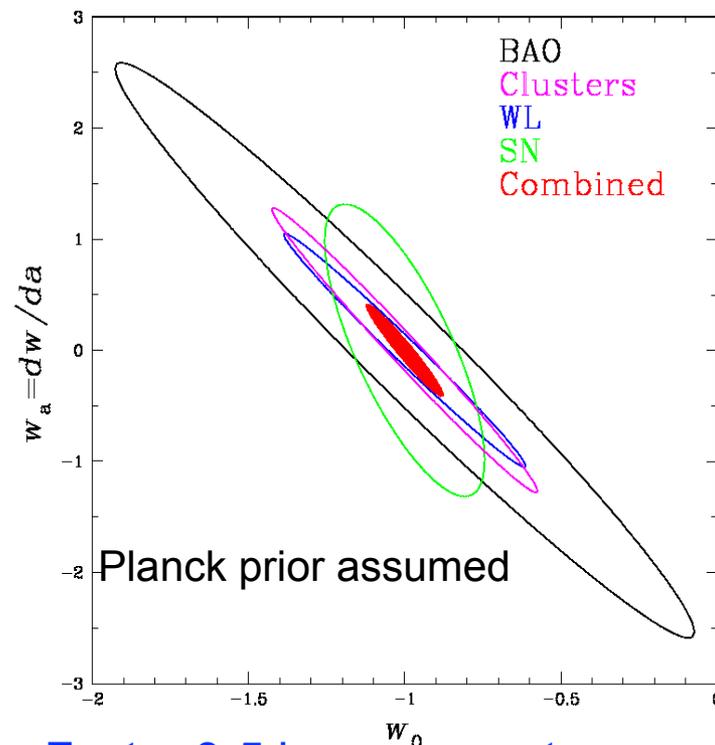
- Shape measurements of 200 million galaxies
- Sensitive to growth of structure and geometry

### • Large-scale Structure (BAO)

- 300 million galaxies to  $z = 1$  and beyond
- Sensitive to geometry

### • Supernovae

- 30-square-degree time-domain survey
- ~4000 well sampled Type Ia to  $z \sim 1$
- Sensitive to geometry



Factor 3-5 improvement over Stage II DETF Figure of Merit



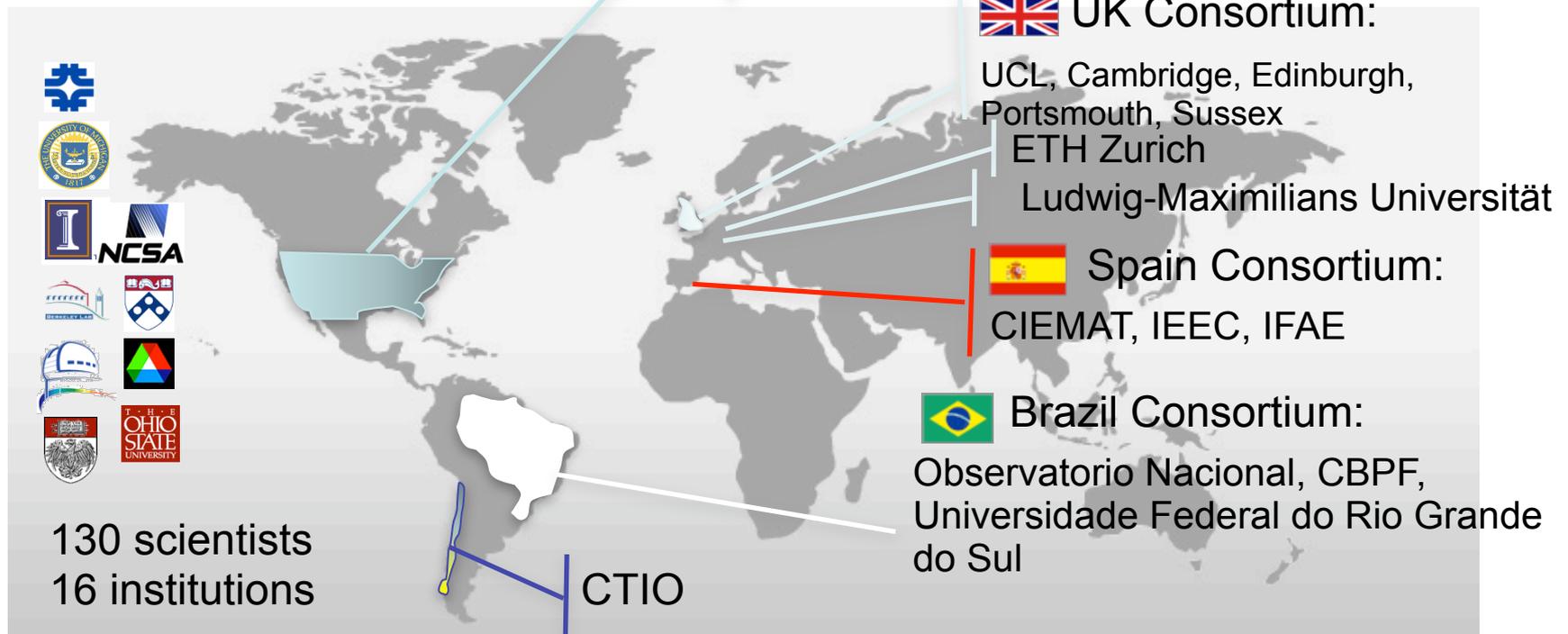
# DES Collaboration

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The DES is an international project to measure cosmic acceleration.

Funding from DOE, NSF, and collaborating institutions.

Fermilab, UIUC/NCSA, University of Chicago, LBNL, NOAO, University of Michigan, University of Pennsylvania, Argonne National Laboratory, Ohio State University, Santa-Cruz/SLAC Consortium, Texas A&M





# Cerro Tololo Interamerican Observatory

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- Blanco Telescope primary mirror support upgraded
- New clean room & control room
- New Telescope Control System
- New Data Transport System





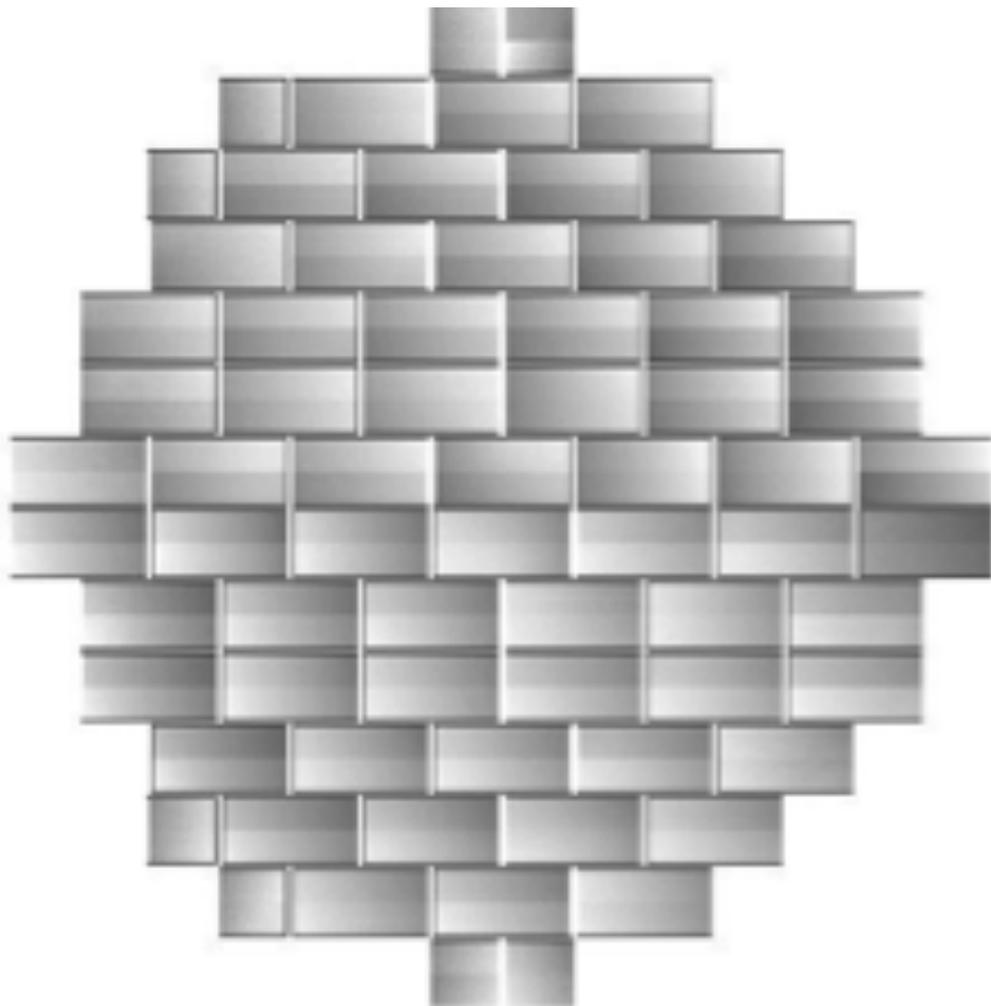
# Data Management

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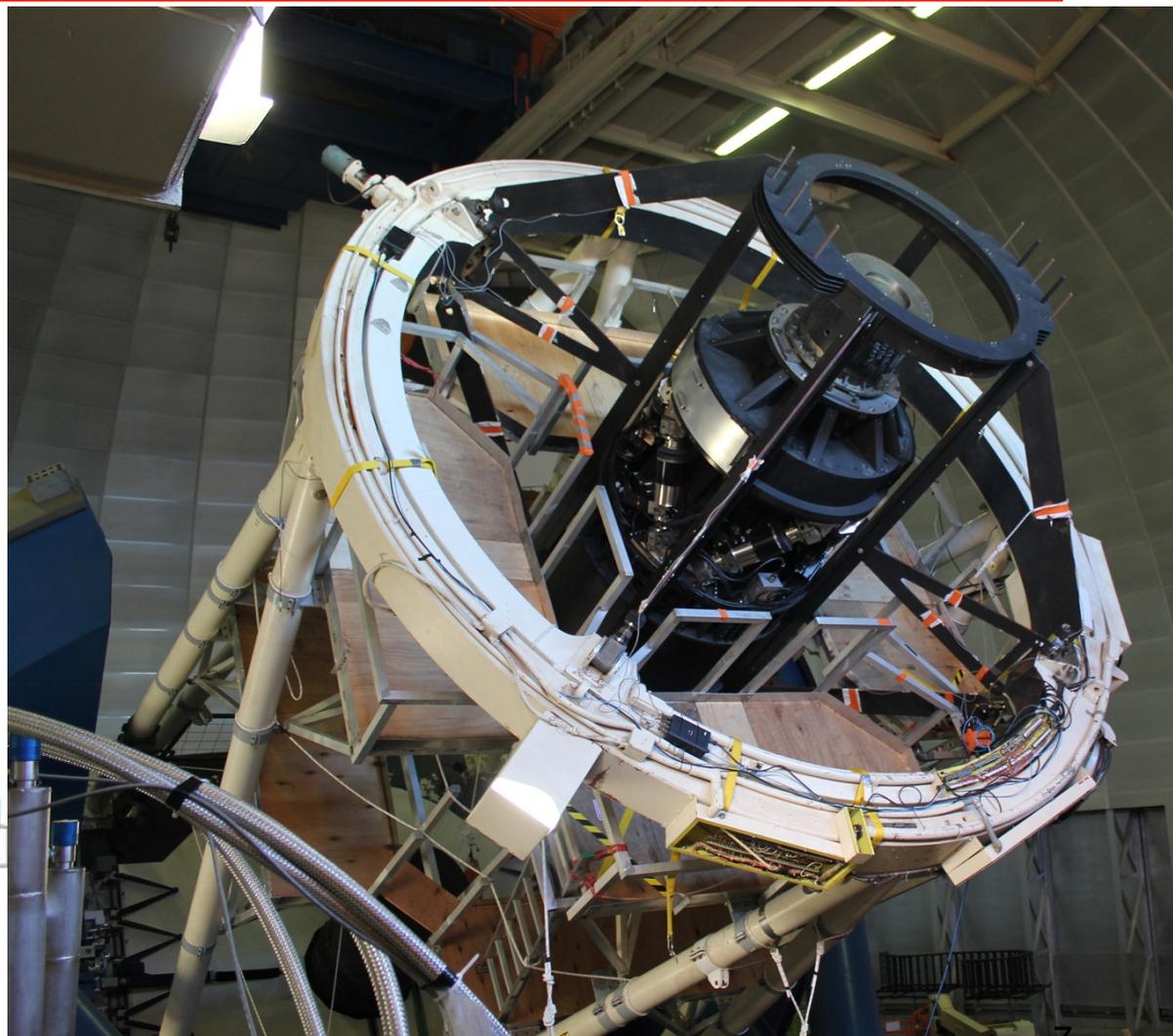
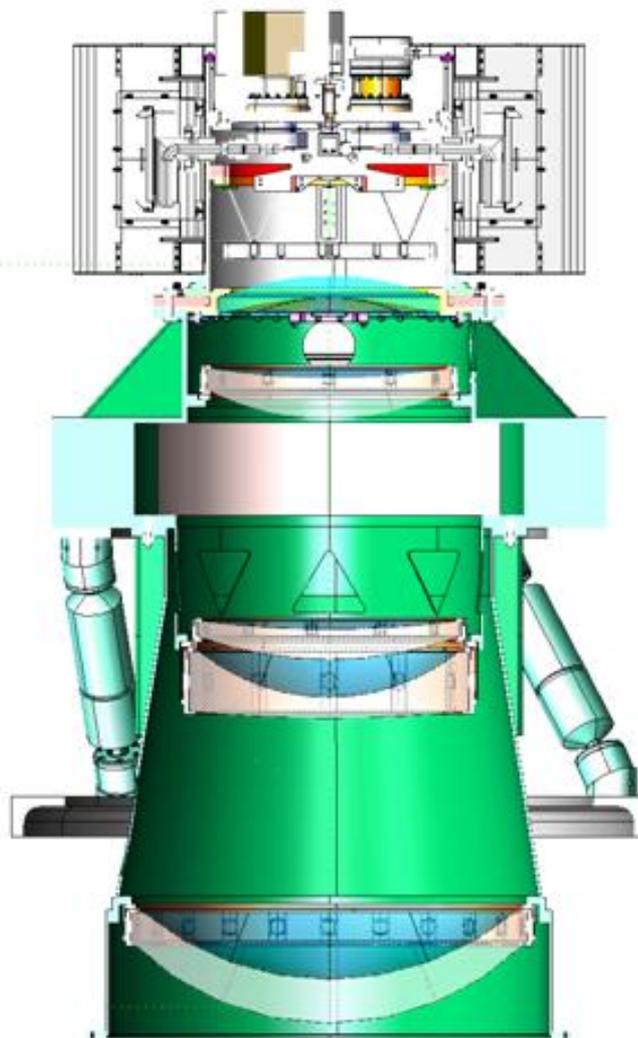
- Each night produces 300 Gby:
  - sent to National Center for Supercomputer Applications for processing
    - galaxy colors (*photometric redshifts*)
    - galaxy shapes (*weak lensing*)
    - identification of Type Ia supernovae (*accurate fluxes*)
- Pipelines tested using simulated data & Data Challenges
- Collaboration actively engaged in data quality testing and analysis, code development and testing





# Dark Energy Camera (DECam)

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# DECam CCDs

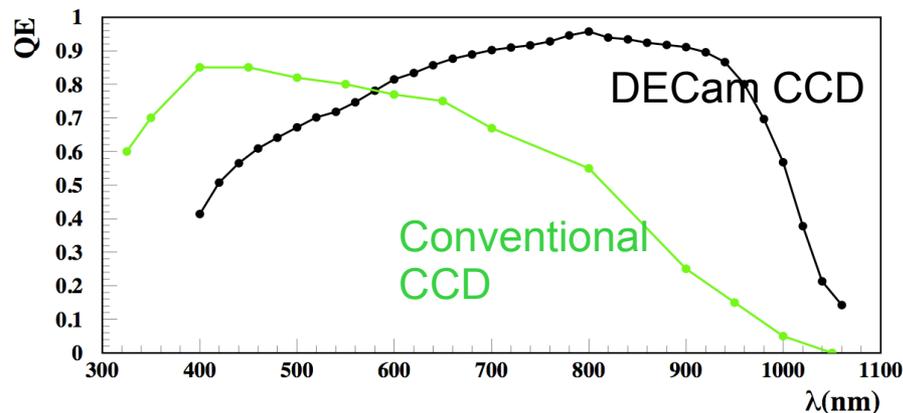
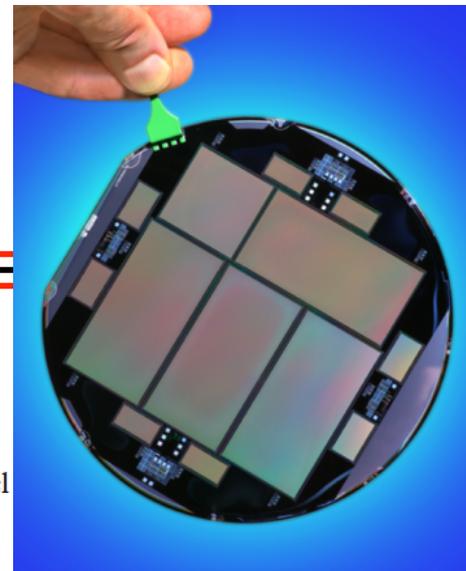
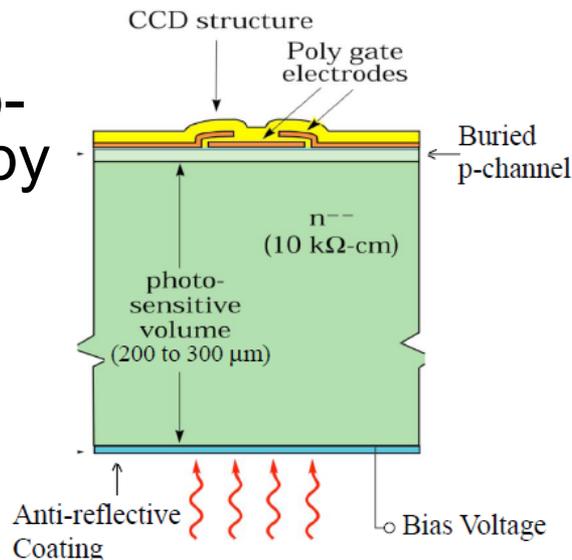
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- The CCDs are thick, back-illuminated, fully-depleted p-channel devices designed by LBNL and manufactured at Dalsa+LBNL

- thinned to 250 microns for DECam
- higher QE over broader wavelength range

- ~500 packaged and tested at Fermilab

- ~120 science-grade devices, 62 in the imager (15-micron pixels  $\Rightarrow$  0.27 seconds of arc)
- leverages FNAL silicon vertex detector construction for the collider programs

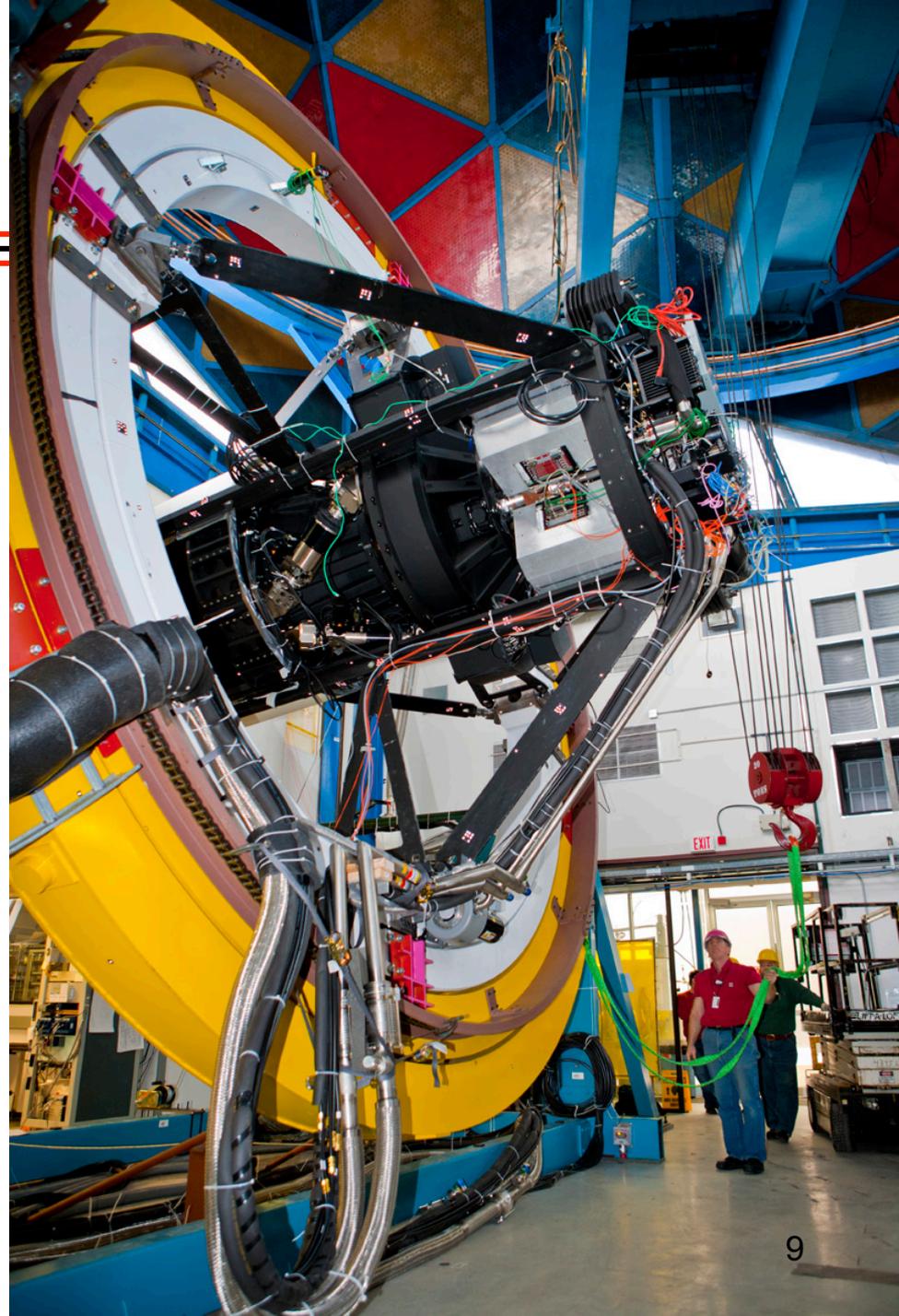




# 2011

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- Telescope Simulator at Fermilab: DECam systems integration
- Essential for testing and debugging all systems prior to shipping; critical for DOE technical reviews
- Devise and practice procedures for actual installation on the Blanco Telescope





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# DECam at CTIO





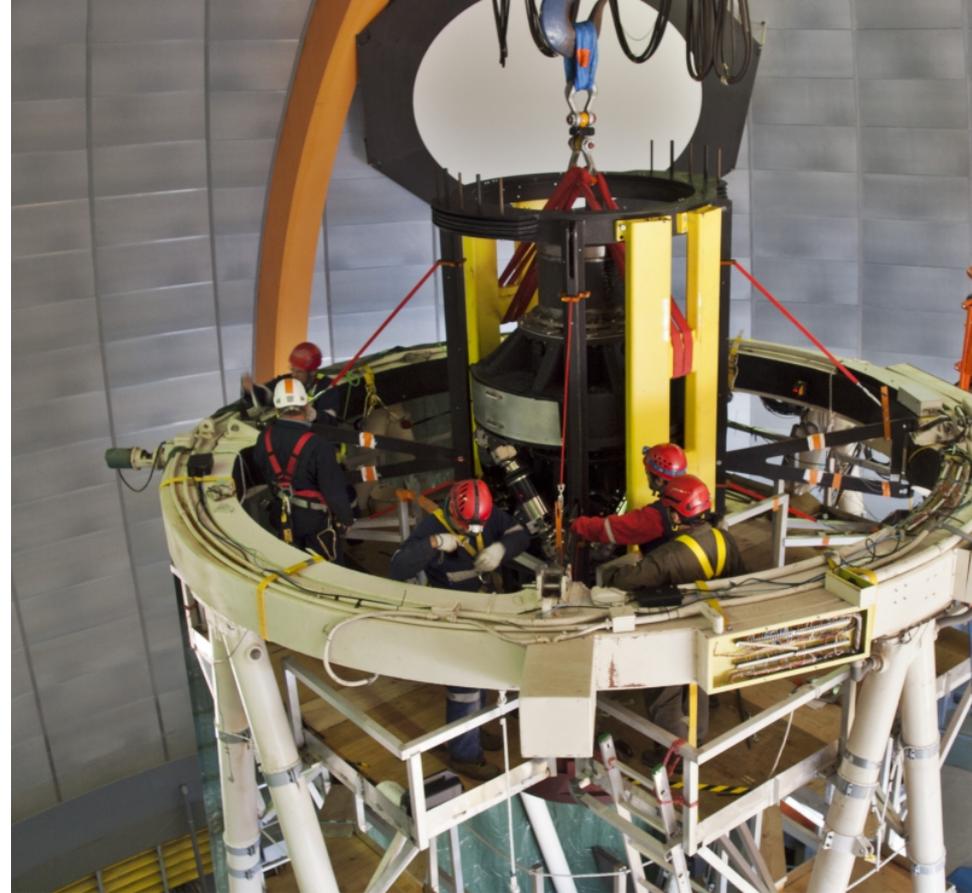
# 2012

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- Old cage removed, new installed (April/May)
- Primary mirror, Cassegrain cage with additional counterweights reinstalled (June)
- Cabling and cryogenics lines (July/August)
- Installation of imager (tomorrow!)
- First light for the imager on the telescope (September)
- Commissioning (Sept. - October)
- Science Verification (November)
- Start of Survey Operations (Dec.)





# DES Collaboration Scientists

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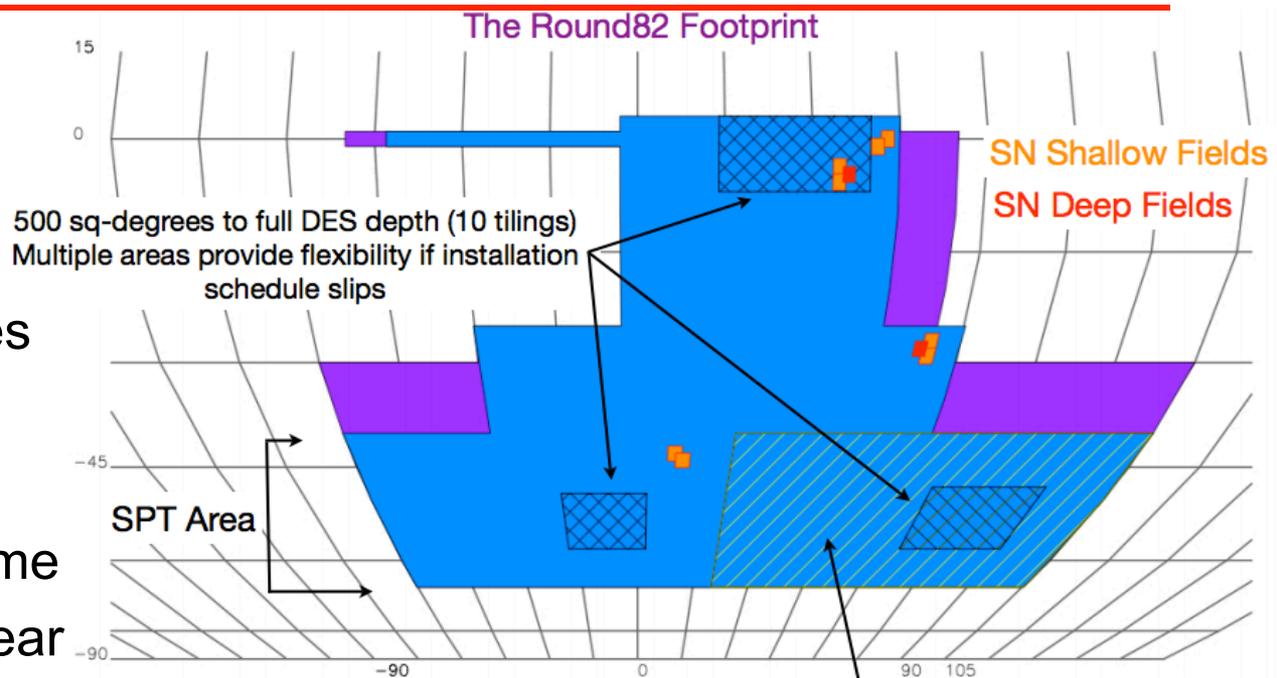
- Finalized survey strategy and evaluated contingency options for first season
- Evaluated Data Challenge 6B quality vs Science Requirements
- Organizing first-season science analysis projects
- Developing spectroscopic follow-up plans for the near and long terms
- Involvement in DECam commissioning and leading DES Science Verification (verify the system is producing survey-quality data prior to operations)



# Survey Strategy

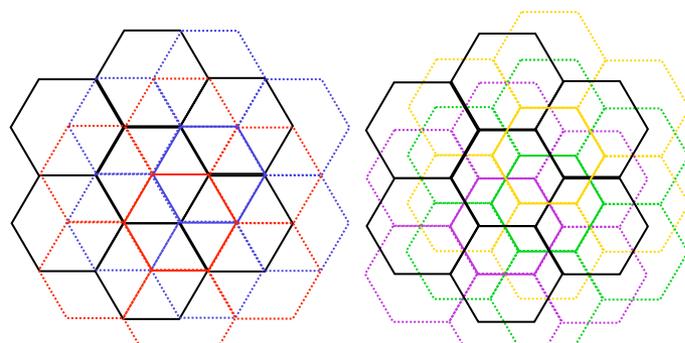
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- 80 - 100 sec exposures
- 2 filters per pointing
  - *gr* in dark time
  - *izy* in bright/grey time
- 2 survey tilings/filter/year
- Interleave 10 SN fields in *griz* if poor seeing or time gap (aim for ~5 day cadence)



2 tilings

3 tilings





# Summary

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- DECam is complete, on budget, and on schedule
- Data Management will be ready for first on-sky data; architecture is being revamped for better operational efficiency, will be phased in after first season
- Successful DOE-NSF partnership, with US and foreign institutional support
- Dark Energy Survey will use DECam to take the next step in constraining Dark Energy
- NSF Astronomy Portfolio Review supports DES and community access to DECam+Blanco